

Flame Spectra at High Temperatures. Part I. Oxy-Hydrogen Blow-Pipe Spectra

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V. *Flame Spectra at High Temperatures.*—Part I. *Oxy-hydrogen Blow-pipe Spectra.*By W. N. HARTLEY, *F.R.S.*

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[PLATES 6, 7.]

SIR DAVID BREWSTER, in 1842, appears to have been the first to examine the spectra of salts by means of the oxygen and coal-gas flame, about 180 of which were deflagrated in a platinum spoon (Edinburgh ‘Roy. Soc. Proc.’ vol. 6, p. 145).

Professor NORMAN LOCKYER* has given us a map of metallic spectra at the temperature of the oxygen and coal-gas blow-pipe (‘Roy. Soc. Proc.’ vol. 23, p. 120). The region observed in the case of twenty-two metals does not extend beyond wave-length 4000 ; and, although we have both arc and spark spectra for the region up to wave-length 1800, we are still unacquainted with the spectra of elements and compounds obtained by means of flames at high temperatures in the ultra-violet region.

Methods of Obtaining Spectra with Flames at High Temperatures.

In studying the spectra of flames there are many points worthy of consideration arising from the structure of the flame, the nature of the combustible, the heat evolved during combustion, and the temperature attained. The temperature of a candle-flame is high enough to give all the spectra capable of being produced by the oxy-hydrogen blow-pipe, for by such simple means we can melt WOLLASTON’S platinum wires and produce the band spectrum of carbon. The reason for such a flame being practically useless for spectroscopic purposes does not arise from the temperature being too low, but from the area of maximum temperature being too small, so that the material to be tested and the support upon which it is held in the flame exercise

* [The following quotation contains a passage which is perhaps the earliest reference to such spectra:—

“The pure earths, when violently heated, as has recently been practised by Lieutenant DRUMMOND, by directing on small spheres of them the flames of several spirit lamps urged by oxygen gas, yield from their surfaces lights of extraordinary splendour, which, when examined by prismatic analysis, are found to possess the peculiar definite rays in excess which characterize the tints of flames coloured by them; so that there can be no doubt that these tints arise from the molecules of the colouring matter reduced to vapour, and held in a state of violent ignition.” ‘Light,’ Sir. J. F. W. HERSCHEL, London, 1827, also ‘Encyclopædia Metropolitana.’ p. 438, vol. 4, 1845.—W. N. H., September 29, 1893.]

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too great a cooling power. A candle or gas flame owes its shape to the rapid ascension of heated combustible vapour and air, or air and gas mixed, and the maximum temperature is to be found near the tip of the flame. The cross section of the flame near its tip should therefore be sufficiently large to completely envelop the support and substance upon it; hence it will be seen that to have a support as small as possible is a distinct practical advantage. For some time a difficulty presented itself in the study of flame spectra of solid substances at high temperatures owing to the necessity which arises for providing an infusible material suitable as a support for the substance to be tested, capable of withstanding the temperature of the oxy-hydrogen blow-pipe flame, and incapable of chemical action upon metallic oxides. I formerly used strips of iridium for the alkaline earths and their salts, but they are quite unsuitable for use with several substances.

I propose to place on record a most convenient method of observing spectra with the oxy-hydrogen flame, and to describe a considerable number of spectra which were photographed preparatory to undertaking the study of spectroscopic phenomena connected with the Bessemer "blow" and the manufacture of steel generally.

The flame of hydrogen, proceeding from a large lead generator, is burnt with compressed oxygen in a small Bunsen blow-pipe, so fixed that the flame is vertical.

The blow-pipe measures 3 inches in length and $\frac{3}{8}$ ths of an inch in external diameter. The substances to be examined are supported in the flame on small plates of kyanite about 2 inches in length, $\frac{1}{20}$ th of an inch in thickness, and $\frac{1}{4}$ th of an inch in width.

This mineral, which is found in large masses in C^o. Donegal, contains 96 per cent. of aluminium silicate, a practically infusible material. It was analyzed in my laboratory some years ago, and owing to the intractable nature of the mineral, the analysis was made with some difficulty.

It is exceedingly difficult to pulverize it, but it readily splits into laminae.

The Instruments and Method of Photography Employed.

The instrument used for the first series of experiments had but one quartz prism of 60°, composed of right and left handed halves, each of 30°. The photographic plates used were "Ilford rapid" and EDWARDS' Isochromatic Plates.

A number of experiments were made with various sensitizers, such as erythrosine used by WATERHOUSE and by MALLMANN and SKOLICK, and cyanine, employed by V. SCHUMANN. Their use proved advantageous in rendering gelatine emulsion plates sensitive to the yellow and red rays.

It was found that diphenylamine blue, used in a similar manner as, and mixed with, cyanine, rendered gelatine-bromide plates rather more sensitive in the region between E and F of the solar spectrum. SCHUMANN has found that emulsions made with 5 parts of silver iodide, precipitated along with 95 parts of silver bromide, are also sensitive in this part of the spectrum.

A trial was made with various developers in order to ascertain which were the most suitable. The spark spectrum of cadmium was photographed on plates of the same kind, with an exposure of five seconds in each case, and development was carefully timed. Developers containing the following reducing substances were used:— (1) pyrogallol, (2) eikonogen, (3) amidol, (4) rodinol, (5) hydroxylamine hydrochloride, (5) hydroquinol, (6) ferrous oxalate, already prepared from potassium oxalate and ferrous sulphate, (7) ferrous oxalate, prepared just prior to use by mixing ferrous sulphate and potassium oxalate solutions kept separate.

Some years ago a similar trial of the then existing developers was made by me and preference was given to hydroxylamine hydrochloride, as prescribed by EGLI and SPILLER, because it gave a brown deposit of silver showing under the microscope no structure or granulation. A commercial sample of the salt, recently purchased, which proved to be strongly acid, was recrystallized from hot alcohol and rendered neutral. It gave good results, but the image was slow in appearing.

Freshly prepared ferrous oxalate was excellent, but best of all was hydroquinol, because it not only produced a dense black image with as much freedom from granulation as any other substance, but it also reproduced lines of feeble intensity, and it developed completely in three minutes as against seven minutes for hydroxylamine, and four or five minutes for other substances.

Granulation appears to be caused by a condition of the gelatine now generally used rather than by the nature of the developing solution as was formerly the case. It was decided to use sensitized plates and hydroquinol as a developer.

Method of Measuring the Positions and Wave-lengths of Lines.

The most convenient and simple method of measuring the spectra emitted by flames is to take a photograph of the spark spectra of tin-cadmium and lead-cadmium alloys superposed upon the former. From the lines of these metals and those of air which accompany them we obtain measurements from which, by an interpolation curve, the oscillation-frequencies and their corresponding wave-lengths may be ascertained.*

The measurement of the lines is made in the same manner as the measurement of the bands in absorption spectra, namely, by simply applying to the photograph an ivory scale which is divided into hundredths of an inch, and by means of a lens or low-power microscope with cross wires in the eye-piece, reading by judgment to tenths of each division. To do this with the greatest accuracy it is necessary to have a straight line ruled down the middle of each spectrum, against which the edge of the scale is fixed in position. To rule this line a very slight nick is made in the jaws of the slit of the spectroscope, which admits more light at this than at any other point, and causes a feeble continuous spectrum to be photographed; upon this the

* Several prominent iron lines beyond $\lambda 3900$ were used in drawing the curve.

lines due to the flame spectra are marked out by the appearance of minute dots. Where the insensitive portion of the film occurs, strong lines are easily seen on the continuous linear spectrum in consequence of the slit being slightly widened for a minute portion of its length, so that the effect caused by want of sensitiveness in the silver salts is diminished.

It is a little difficult to read the measurements and describe the spectra at the same time, hence enlargements were made upon which the measurements were recorded as they were read off. Another convenient plan was to adjust the scale to the photograph and take an enlargement therefrom at once, so that prints from the same give approximately their own measurements. Only those measurements are exact which are exactly at the centre of the photographic lens, even when the scale is precisely adjusted to the photograph, so, for instance, that the 20th division was exactly at the sodium line. In cases where the lines were not newly discovered, and it was only necessary to identify them, nothing more was required. New lines and bands were measured by a micrometer screw with a pitch of 100 threads to the inch, and a wheel head divided into 100 parts. The screw carries a nut on which a microscope, magnifying 10 diameters, is fixed, by which arrangement it is easy to measure to $\frac{1}{10,000}$ th of an inch, and, where desirable, to $\frac{1}{100,000}$ th. This instrument was made by Mr. A. HILGER, of London. Each measurement was recorded at the time by writing on an enlarged print of the same photograph.*

The Spectrum seen when supports of Kyanite alone are heated in the Oxy-hydrogen Flame.

Just as in the ordinary use of the spectroscope we are prepared to see the lines of sodium, and under certain circumstances the bands peculiar to carbon, so in these photographs, the sodium lines and the strongest groups of lines belonging to the emission spectrum of water vapour, are also always present. In addition to these, the kyanite yields the red line of lithium, which is no inconvenience, but a positive advantage, as it serves to indicate where the spectrum commences, and from which point measurements may be made.

The Extent and Character of the Spectra observed.

Although the apparatus is capable of photographing on one plate rays lying between wave-lengths 6708 of lithium in the red and 2194 in the ultra-violet, nevertheless the flame spectra of a large majority of the metals and their compounds terminate somewhere about the ultra-violet emission spectrum of water. The first,

* For several of the enlarged negatives made exactly to the same scale I am indebted to the kindness of my friend, Professor ALEC FRASER, who devoted much of his own valuable time to making negatives with as perfect a definition as possible, the prints from which have greatly facilitated my work.

second, and third series of lines measured, LIVEING and DEWAR, always appear in these spectra; in some cases the fourth and fifth series are well rendered.

Although the number of lines exhibited by some of the metals is large, yet the extent of spectrum is small compared with that yielded by condensed sparks. Typical band spectra are exhibited by sulphur, selenium, and tellurium. The first yields a continuous spectrum, in which a series of beautiful bands is seen, the second a series of fine bands occurring at closer intervals, the third is characterized by bands still closer together, and near the more refrangible termination of which four lines occurring in the spark spectrum of tellurium are visible.

Thus we see that increase in atomic mass causes shorter periods of recurrence of bands, while we know that it causes greater periods in the recurrence of lines.

It is also worthy of remark that the most volatile of these elements emits a continuous spectrum, with a band spectrum just emerging from it; the second gives us a beautiful and purely a band spectrum, while the third least volatile and more metallic substance of largest atomic mass, and producing the densest vapour, yields a band spectrum, together with a line spectrum. Several metals, such as nickel, yield nothing but lines, others give us both lines and bands, as manganese and iron, while tin, lead, silver, and gold yield very beautiful band spectra. Metalloids and non-metallic elements are generally considered to be essentially different from metals, since they emit channelled or band spectra at one temperature and line spectra at another. It was, in fact, first laid down by PLÜCKER and HITTORF that "*There is a certain number of elementary substances, which, when differently heated, furnish two kinds of spectra of quite a different character, not having any line or any band in common*" ('Phil. Trans.,' vol. 155, p. 6).

The discovery of this fact was of great importance, for it led to the conclusion that as one spectrum of an element is replaced by another and totally different spectrum of the same element, there must be an analogous change in the constitution of the ether, indicating a new arrangement of the gaseous molecules, and this implies either a chemical decomposition, or an allotropic condition of the vapour of the substance. PLÜCKER and HITTORF concluded that the same matter, in two allotropic states, emitted different spectra, but the allotropy was dependent solely on temperature. Band spectra they designated spectra of the 1st Order, and Line spectra, spectra of the 2nd Order. The former have been fully recognized as the spectra of metalloids, such as carbon, phosphorus, sulphur, selenium, and tellurium, but it seems to have been overlooked that PLÜCKER and HITTORF observed spectra of the 1st Order in the case of a few heavy metals, particularly lead and manganese. Metallic lead and its compounds were found to yield the same band spectrum in the oxy-hydrogen flame, and manganese exhibited a curious spectrum of the 1st Order, most similar to that of carbon, but with the lines composing the bands differently distributed. The well-known spark spectra of these elements are spectra of the 2nd Order.

LECOQ DE BOISBAUDRAN has observed a beautiful spectrum of aluminium of the

1st Order, obtained by means of an uncondensed spark. That this metal at so high a temperature yields such a spectrum is undoubtedly due to the fact that it is almost, if not absolutely, impossible to vaporize it with the oxy-hydrogen flame.*

LIVEING and DEWAR have recently obtained a band spectrum by the combustion of nickel tetracarbonyl which is also accompanied by lines ('Roy. Soc. Proc.,' vol. 52, p. 117). This spectrum, I expect, will be found to be due to metallic nickel and not to the compound substance.

Yttrium and scandium, in solutions of their chlorides, each yield a line spectrum, with a group of bands in the red and orange region, when submitted to the action of a condensed spark. From the foregoing facts, and from the descriptions of spectra which here follow, it will be seen that several metallic elements emit banded spectra.

Characteristic Flame Spectra of Elements emitted at High Temperatures.

- I. *Line Spectra.*—Lithium, thallium, nickel, cobalt.
- II. *Band Spectra.*—Antimony, bismuth, gold, tin, sulphur, selenium.
- III. *Band Spectra with Lines.*—Copper, iron, manganese, tellurium, lead, and silver.
- IV. *More or less continuous Spectra with Lines.*—Sodium, potassium, magnesium, chromium, cadmium.
- V. *A continuous Spectrum.*—Zinc, carbon, arsenic, aluminium.
- VI. *No Spectrum.*—Platinum.

It might be supposed that the band spectra were due to the oxides and not to the metallic elements in Group II., but there is evidence against this in the case of silver† and gold, since no oxides of these metals can exist at the temperature of the flame employed.

In the case of manganese the evidence is of a different character, and may be referred to at somewhat greater length, since MARSHALL WATTS has attributed the band spectrum seen in the Bessemer flame to the oxide of manganese, chiefly on the ground that it was yielded by manganese chloride (*Spectres Lumineux*), and in the oxy-hydrogen flame by manganic oxide. No evidence was adduced to show that the spectrum in either instance was due to the metal.

* See Appendix 5, p. 211.

† [Channelled emission spectra of silver and tin, produced by the electric arc, have been noticed by LIVEING and DEWAR.

"Tin gives flutings in highly refrangible portions of the spectrum, and silver gives a fine fluted looking spectrum in the blue." 'Roy. Soc. Proc.,' vol. 34, p. 122, 1882.

The same observers have described the channelled spectrum of magnesium oxide. A set of seven bands in the green beginning λ 5006-4 and fading towards the violet side of the spectrum are stated to be due to the oxide or to the process of oxidation. 'Roy. Soc. Proc.,' vol. 44, p. 243.—W. N. H. September 29, 1893.]

On the other hand, the evidence that it is due to the metal is of the following character:—

(1.) It may be produced from the metal in a reducing flame, and it disappears when an excess of oxygen is present. (2.) Although it may be produced by heating manganic oxide containing 66 per cent. of manganese, the spectrum is weak. (3.) A stronger spectrum is obtained by heating spiegel-eisen containing 18 to 20 per cent. of manganese, and by heating ferro-manganese, containing 80 per cent. of manganese, than that which it is possible to obtain by heating, to the same temperature and during the same period, manganic oxide containing 66 per cent. of manganese. Silico-spiegel containing 10 per cent. of silicon and 18 to 20 per cent. of manganese did not yield the manganese bands so strongly as the spiegel-eisen containing the same proportion of metal, probably because the manganese is converted into silicate. Even TURTON'S tool steel yields a fairly strong indication of the manganese bands.

If we examine the spectrum of air of the first order as obtained by sparks uncondensed, it appears to consist of bands only, but a more minute examination of spectra taken with an instrument giving considerable dispersion and excellent definition has shown that the bands are composed of three over-lapping series of lines. Such a character is usual with degraded band spectra of elements. If the pressure be reduced from the normal of 760 millims. to something like 5 millims. or less, then the bands disappear, and the strongest edge of each band remains as a line to represent the spectrum of the element at diminished pressure. Now, this change is one which is observed in the case of those metals which give band spectra, but, if they give bands and lines together, then the lines remain after the bands have vanished. This is to be observed in the spectra of silver, lead, bismuth, and tellurium.

The most interesting case, however, is that of silver, for the spectrum is composed of a number of regularly disposed and closely placed lines.

The bands are degraded towards the rays of lesser refrangibility, that is to say, in this direction the lines are of diminishing intensity, and they are of increasing width apart. When the quantity of silver diminishes, and consequently the vapour exerts less pressure, being mixed with the vapour of other metals, the bands become narrower until at last nothing but lines remain, and these are the strongest lines belonging to the strongest bands. They correspond to those on the spark spectrum with wavelengths 3382·3 and 3280·1.

Thus we see how the line spectra are related to band spectra, and that there is really no essential difference between the constitution of the matter which enters into the vapours of metals and metalloids; there is, in fact, something in their constitution common to both, which is apparently dependent on their vapour pressures and probably due to the action of the molecules upon one another when

their mean path is so extended that their motions become rhythmical. Reduce the freedom of their motions and the result is a continuous spectrum.*

MITSCHERLICH first drew attention to the distinct spectra, for the most part composed of bands, which are emitted by compounds ('Pogg. Annalen,' vol. 121, p. 459).

DIACON also ('Thèses de Physique et de Chimie,' Montpellier, 1864, Boehm et fils), using a flame the interior of which was fed with chlorine, obtained distinct spectra of chlorides such as those of the alkaline earths, also gold, lead, iron, cobalt, and nickel.

The alkalis gave no spectrum except where the conditions were such that they became converted into oxide or metal, as in the mantle of the flame. Of the various compounds examined, some gave degraded band spectra, others plain bands, and many yielded line spectra, or bands and lines together. PLÜCKER and HITTORF first showed that the alkali metals and their salts emit, even at a low temperature, spectra of the 2nd Order or lines, while metals of the alkaline earths, and compounds of the same emit band spectra, accompanied by a principal line. When the bands are well developed they constitute a spectrum of the 1st Order; this was proved in every respect to be the case with the band spectrum of barium.

Flame Spectra Emitted by Compounds at High Temperatures.

I. *Spectra of Elements. Chiefly Lines.*—Iron, nickel, cobalt, chromium, manganese, sodium, potassium, lithium, thallium, rubidium.

II. *Spectra Peculiar to Compounds. Lines and Bands together.*—Calcium oxide and salts, calcium fluoride, strontium oxide and salts, barium oxide and salts, beryllium oxide and salts, magnesium oxide and salts, aluminium oxide and salts, cadmium oxide and salts, copper oxide and salts, chromic trioxide, phosphorus pentoxide, cerium oxide and salts, cerium chloride.

The study of the spectra of compounds is one of much interest, particularly in its bearings on the periodic law, and the prosecution of this subject is being continued.

Application of the Oxy-hydrogen Flame Spectra to Chemical Analysis.

Alkali Metals.—The examination of insoluble minerals, such as silicates, in order to detect the alkali metals, may be readily made with the oxy-hydrogen blow-pipe. Proof of the presence of lithium and sodium in kyanite is evidence of this. My assistant, Mr. RAMAGE, examined a microcline felspar from the granite of Dalkey, Co. Dublin, by fixing a piece of it in the flame for half-an-hour while a photograph was taken. The lines of sodium, potassium, lithium, and rubidium were identified.

Alkaline Earth Metals.—A piece of dolomite gave the lines and bands characteristic

* See Professor SCHUSTER'S British Association Report, 1880.

of lime with the bands of magnesium. The sulphates of calcium, strontium, and barium readily yield their spectra by exposure to the flame.

Metals Yielding Band Spectra.—These are elements of considerable volatility, the lines of which become converted into bands as their proportion in the substances to be examined diminishes.

The lines which serve for the detection of small quantities of the respective elements are the following :—

	λ .	
Copper	3273·2 3246·9	
Silver	3382·3 3280·1	
Tin	3033·1 3007·9	
Lead	4059 3684 3639·5 2832·2	Mean of $\left\{ \begin{array}{l} 4061·5 \\ 4057·6 \end{array} \right\}$ Or (3682·9) (3639·2)
Thallium	(5349·6) 3775·6	
Bismuth	4724·5 3067·0	Approximately
Cadmium	(3261·17)	
Manganese bands	$\left\{ \begin{array}{l} 5845 \text{ to } 5700 \\ 5700 \text{ ,, } 5645 \\ 5645 \text{ ,, } 5591 \end{array} \right.$	
„ lines	(4031·8) (4029·9)	These lines are visible after the bands have disappeared most completely

As an illustration of the method of testing for these substances it may be mentioned that a finely crystallized specimen of bismuth was volatilized in the flame. A number of conspicuous lines on the photograph were measured with the ivory scale and their wave-lengths ascertained. Thus two lines were identified with thallium, three with lead, two with copper, two with silver, and the remainder proved to be bismuth lines. Copper was detected in steel.

Metals which emit Line Spectra.—The spectra of these elements are somewhat complicated, taking for instance iron, nickel, and cobalt, as examples. Iron is readily

detected by the groups of lines lying between 3929·7 and 3749·4, also between 3745·4 and 3683. Chromium is recognized by its two sets of triplets. A more particular examination of nickel and cobalt has not yet been made with a view of ascertaining their most persistent lines.

The prominent manganese lines were detected in the spectra obtained from malleable cobalt and nickel, also in fine steel.

Descriptions of Spectra and Measurements of Lines and Bands, with their approximate Wave-lengths.

THE OXY-COAL-GAS FLAME.

The flame was non-luminous. Photographs were taken with a somewhat wide slit, and the exposure was one hour. The edges of the bands are as sharp as they are generally seen in the spectrum of a Bunsen flame, and the lines of which the bands are composed are somewhat wide. No attempt was made to purify the coal-gas, as the object of examining this spectrum was to determine the origin of any lines which might be caused by hydrocarbons in the oxy-hydrogen flame. LECOQ DE BOISBAUDRAN has carefully described variations in the spectrum seen under different circumstances in the flame of a Bunsen burner, but there is no occasion to refer to these further. (*Spectres Lumineux.*)

All the principal bands observed are probably due partly to carbon and partly to what is generally considered as the cyanogen spectrum. They are indicated by (C) carbon, and (CN)₂ cyanogen. The measurements of lines and bands made by other authors are indicated thus:—K. and R., KAYSER and RUNGE; L., LECOQ DE BOISBAUDRAN; L. and D., LIVEING and DEWAR; D., DESLANDRES; W., WATTS; F., FIEVEZ. The lines and bands were all measured twice and their wave-lengths ascertained on two separate occasions. β , γ , δ , ϵ are groups or bands described by LECOQ DE BOISBAUDRAN.

THE OXY-COAL-GAS FLAME.

Micrometer measurements in hundredths of an inch.	Description of the spectrum.	$\frac{1}{\lambda}$.	λ .	Remarks.
26.02	The fainter edge of the 1st band	17671.5	5659	5627.5 (C), F. 5610.8 (C), F. 5577.7 } (C), F. 5577.4 } 5577.0 } 5557 (C), F. This is the δ group described by LECOQ DE BOISBAUDRAN. Yellow rays. 5478.4 (C), W. 5440 (C), W. 5425 (C), W. } } 5165.5 (C), W. } 5138.56 } (C), K. and R.; } 5138.84 } 5138.4, F. } 5138.13 } } 5097.9, F.; 5100, W. } 5098.84 } (C), K. and R. } 5098.19 } } 5086.9 (C), F.; 5082 (C), W. } 5086.43 } (C), K. and R. } 5086.31 } 4951.5 (C), K. and R. 4899.98 (C), K. and R. 4815.66 (C), K. and R. 4775.82 (C), K. and R. 4763.86 (C), K. and R.
26.92	" stronger "	17775.5	5627	
27.32	The fainter edge of the 2nd band	17821	5611	
28.33	" stronger " "	17938	5577	
28.82	The fainter edge of the 3rd band	17993.6	5557	
29.93	" stronger " "	18119.2	5520	
30.72	The stronger edge of the 4th band overlapped by the foregoing δ	18207	5492	
31.31	" " " " "	18272	5473	
32.1	" " " " "	18360.5	5446	
32.88	" " " " "	18447	5422	
33.54	" " " " "	18521	5399	
34.38	" " " " "	18614	5372	
40.45	The less refrangible edge of the 10th band	19257.5	5193	
41.21	" more " " "	19341.5	5170	
42.32	A marking like the darker edge of the 11th band	19463.2	5138	
43.68	" " " " " 12th "	19613	5098	
47.0	" " " " " 13th "	19962	5086	
49.2	The darker edge of a faint band overlapped by foregoing (14th)	20192.2	4952	
51.35	The less refrangible edge of a broad band (15th)	20415	4899	
54.7	Marking in of a broad band	20762.5	4816	
56.47	The more refrangible edge of the same	20946.5	4774	
56.89	A band is overlapped by this.	20989.5	4765	

THE OXY-COAL-GAS FLAME—(continued).

Micrometer measurements in hundredths of an inch.	Description of the spectrum.	$\frac{1}{\lambda}$	λ .	Remarks.
57-765	<p>This is the stronger portion of the γ group described by LECOQ DE BOISBAUDRAN. Markings like sharp lines in the 16th band</p> <p>The stronger or more refrangible edge of the same</p> <p>This band also overlaps another diffused band.</p>	21081	4743.5	<p>4739.8 (C), W. 4731.9 (C), F. 4732.33 (C), K. and R. 4731.93 (C), K. and R. 4719.87 (C), K. and R. 4720.1 (C), F.; 4717.2 (C), W. 4702.3 (C), K. and R. 4702 (C), F.; 4698.4 (C), W. 4688.2 (C), K. and R. 4688.9 (C), F.; 4684.2 (C), W.</p>
58-24		21130.2	4732	
58-79		21187	4720	
59-56		21267	4702	
60-22		21335	4688	
60-56	<p>Markings in the 17th band. This is the weaker part of the γ group.</p> <p>Stronger edge of the same</p> <p>The edge of a marking like a broad band or a continuous spectrum, very diffused (18th band)</p> <p>Markings like lines in the 18th band</p>	21370	4679	<p>4678.9 (C), F.; 4677 (C), W. 4672 (C), F.; 4679 (C), W. 4672.2 (C), F.; 4675 (C), L.</p>
60-91		21405	4672	
71-0		22414	4462	
74-0		22709	4405	
74-485 75-49	<p>Markings like lines in the 18th band</p> <p>The stronger and more refrangible edge of the same</p> <p>The stronger edge of a fainter band overlapped by the foregoing</p> <p>A faint though sharp line</p> <p>The sharp, less refrangible edge of a strong broad line merging into a band or continuous spectrum</p> <p>Probably the more refrangible edge of the same line, or may be only a marking in the above band or continuous spectrum.</p> <p>Broad line in the continuous spectrum or band</p>	22754 22847.2	4395 4378	<p>4381 (C), L. and D. About the centre 4368 (C), L. 4365 (C), L. and D.</p>
76-22		22916.8	4364	
76-95		22985	4350	
77-51		23035	4342	
78-085	23086.4	4332	4332	
79-305	23194	4312	4312	4311 (C), W.
79-86	23246	4302	4302	<p>4365 (C), L. and D.</p>
80-68	23322.2	4288	4288	
81-07	23358.5	4282	4282	
81-52	23400	4273	4273	
81-95	23436	4268	4268	
82-3	23472.5	4260	4260	

FLAME SPECTRA AT HIGH TEMPERATURES.

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THE OXY-COAL-GAS FLAME—(continued).

Micrometer measurements in hundredths of an inch.	Description of the spectrum.	$\frac{1}{\lambda}$.	λ .	Remarks.
82·61	Marking in the continuous spectrum, or band	25501	4255	
83·1	" "	23547	4248	
83·56	" "	23589·5	4240	
84·21	" "	23649·5	4230	
85·0	" "	23722	4215	4215·26 (CN) ₂ K. and R.
85·57	" "	23773·5	4208	4208·4 (CN) ₂ K. and R.
86·25	" "	23836	4196	4196·05 (CN) ₃ K. and R.
There were several other lines which were too faint to measure with a magnifying power of 10 diams.				
Then appears a series of beautiful very fine and closely adjacent lines, numbering sixteen in all.				
99·22	First line of the series	24985	4003	
99·6	Second "	25012	3998	
100·05	Third "	25052	3992	
100·64	Fourth "	25104·6	3984	
101·44	Fifth "	25176·5	3973	
102·135	Sixth "	25236·2	3963	
102·88	Seventh "	25298	3954	
103·36	Eighth "	25341	3946	
103·96	Ninth "	25391·3	3938	
104·45	Tenth "	25432·5	3932	
104·9	Eleventh "	25470·4	3926	
105·46	Twelfth "	25516·2	3920	
105·8	Thirteenth "	25545	3913	
106·2	Fourteenth "	25579	3908	
106·61	Fifteenth "	25613·5	3904	
107·06	Sixteenth "	25652	3898	
107·5	Less refrangible edge of a broad line or narrow dark band	25689·4	3893	3893·1 (C) ₂ D.
108·35	More "	25761·5	3882	3883·1 (CN) ₂ D.

THE OXY-COAL-GAS FLAME—(continued).

Micrometer measurements in hundredths of an inch.	Description of the spectrum.	$\frac{1}{\lambda}$	λ .	Remarks.
109.45	Faint line or less refrangible edge of a faint band	25855.5	3868	3871.4 (CN) ₂ , D.
110.48	" " marking in the same	25932.2	3856	3855.06 (CN) ₂ , K. and R. 3854.7 (CN) ₂ , K. and R.
111.17	" " "	26001.5	3846	3839.98 (CN) ₂ , K. and R.
111.63	" " "	26040.4	3840	3831.15 (CN) ₂ , K. and R.
112.32	" " "	26100	3831	3825.4 (CN) ₂ , K. and R.
112.81	" " "	26141.3	3825.5	3825.1 (O), D. 3823.9 (CN) ₂ , K. and R.
113.0	" " "	26157.8	3823	3819.36 (CN) ₂ , K. and R.
113.45	" " "	26190.5	3818.3	3816.24 (CN) ₂ , K. and R.
113.78	Less refrangible edge of a faint band	26214.2	3815	3790
115.8	More " " the same	26386	3642.5	3642.63 (CN) ₂ , K. and R.
128.74	Less refrangible edge of a faint band	27455	3579.5	3579.22 (CN) ₂ , K. and R.
134.81	More " " the same	27943	3568.5	3568.4 (CN) ₂ , K. and R.
135.79		28022	3563	3563.92 (CN) ₂ , K. and R.
136.37	Two faint lines, like markings in or edges of bands	28067	3544.5	3545.07 (CN) ₂ , K. and R.
138.28		28212	3528	3528.71 (CN) ₂ , K. and R.
140.0	A faint line, like a marking in or edge of a band	28343	3522	3522.49 (CN) ₂ , K. and R.
140.65	" " " " " "	28392.8	3498.5	3497.7 (CN) ₂ , K. and R.
143.17	" " " " " "	28584	3487.8	3487.61 (CN) ₂ , K. and R.
144.33	" " " " " "	28672.5	3478.8	
145.29	" " " " " "	28746	3478.8	
145.84	" " " " " "	28788	3473.6	
148.21	" " " " " "	28969	3452	
148.62		29001	3448	
148.91		29022.5	3445	
149.42		29062	3441	
149.79		29090	3437	
153.82		29395	3402	
154.71		29461.5	3394	
155.9		29551	3384	
157.09		29640	3373	
158.86		29771.5	3359	3360.1 (CN) ₂ , D.
160.145		29866	3349	
161.62		29975.5	3336.5	

THE OXY-COAL-GAS FLAME—(continued).

Micrometer measurements in hundredths of an inch.	Description of the spectrum.	$\frac{1}{\lambda}$.	λ .	Remarks.
162.43	A faint line	30035.5	3330	3305.3 (C), D.
163.48	"	30112.5	3321	
164.47	"	30186.5	3313	
165.55	"	30266.5	3304	
166.17	"	30312.5	3299	
167.5	"	30411	3288	
168.72	"	30501	3278	
170.0	"	30595	3269	
171.65	"	30717	3256	
	There are a few more of these lines, which, however, were too faint to measure accurately with a magnifying power of ten diameters.			

THE CARBON MONOXIDE FLAME.

The gas was burnt from a blow-pipe along with oxygen. The plate was exposed for one hour. The spectrum is continuous from about λ 5800 to about λ 3000.

A somewhat wide slit was used as in photographing the oxy-coal-gas flame.

Certain broad lines occur on the continuous rays, and these for the most part have been identified with certain lines occurring in the spectrum of carbon, as the measurements approximate closely to some of those taken from the spectra of this element observed by MARSHALL WATTS, ÅNGSTRÖM and THALÉN, and PIAZZI SMYTH.

The very strong and extended continuous spectrum is a remarkable feature of this, as it is likewise in that of the Bessemer flame spectrum.

Micrometer measurements in hundredths of an inch.	Description of the spectrum.	$\frac{1}{\lambda}$.	λ .	Remarks.
9.7	A faint line or band, marking scarcely visible, not sharp but indistinct.	15782	6337	
13.5	" " " " " " " " " " " "	16208	6172	5953.5 (C), Å. and T.;
18.75	" " " " " " " " " " " "	16822	5945	5955 (C), W.
23.0	An exceedingly faint line	17318	5777	5534.5 (C), Å. and T.
29.5	A faint line	18070	5534	5473 (C), P. S.
31.3	" " " " " " " " " " " "	18271	5473	5165.5 (C), W.
41.3	" " " " " " " " " " " "	19352	5168	5036.7 (C), Å. and T.
46.0	A very faint line or band marking, just barely an indication that there is an inequality in the continuous spectrum. This same description applies to all that follows	19857	5037	
48.5	" " " " " " " " " " " "	20119	4970.5	4969 (C), Å. and T., band; also line, W.
49.5	" " " " " " " " " " " "	20223	4945	4947 (C), line, W.
62.5	" " " " " " " " " " " "	21555.5	4640	4637 (C), line, W.
64.75	" " " " " " " " " " " "	21790	4589	4585 (C), line, W.
71.8	" " " " " " " " " " " "	22494	4446	
83.0	" " " " " " " " " " " "	23537.5	4249	4249 (C), band, W.
84.5	" " " " " " " " " " " "	23676	4224.5	
87.0	" " " " " " " " " " " "	23904.5	4183	

LITHIUM.

Lithium chloride. Exposure 30 minutes. KAYSER and RUNGE's measurements refer to arc spectra of the alkalis and alkaline earths. 'Ueber die Spectren der Elemente,' Königl. Preuss. Akademie, 1890, IV.

Ivory scale numbers.	$\frac{1}{\lambda}$	KAYSER and RUNGE's measurements.		
		$\frac{1}{\lambda}$	λ .	
2.8	..	1490713	6708.2	P.s.
64.1	2173	2172794	4602.37	D.s.
90.3	2420	2419878	4132.44	D.s.
174.55	3094	3093322	3232.77	P.s.

P.s. Principal series.

D.s. Diffuse series.

SODIUM.

Sodium chloride. A perfectly pure specimen specially prepared. Exposure 35 minutes. A very strong continuous spectrum extends from λ 6020 to 3600, it continues weakly to λ 3320. *Loc. cit.*, KAYSER and RUNGE.

Ivory scale numbers.	$\frac{1}{\lambda}$	λ .	KAYSER and RUNGE's measurements.		Remarks.
			$\frac{1}{\lambda}$	λ .	
6.0	15340	6518	Bands and lines not previously observed. Some rather broad, others narrow. Band with lines upon it.
8.0	15574	6420	
9.5	15748	6349	
10.8	15898	6290	
11.2	15946	6271	
12.3	16042	6233	
14.2	16290	6138	Stronger edge of band.
16.8	16595	6026	Stronger edge of band at 15.81.
20.0	16975	..	(1696091)	(5896.16)	Centre of band with stronger edge at 17.25.
			(1697738)	(5890.19)	D ¹ P.s.
25.3	17575	..	(1758007)	(5688.26)	D
			(1759665)	(5682.9)	D.s.
47.8	2007	..	(2006610)	(4983.5)	D.s.
			(2008314)	(4979.3)	D.s.
61.1	2142	..	(2141603)	(4669.4)	D.s.
			(2143531)	(4665.2)	D.s.
165.6	30280	..	(3027487)	(3303.07)	P.s.
			(3028037)	(3302.47)	P.s.

P.s. Principal series.

D.s. Diffuse series.

POTASSIUM.

Potassium chloride. Exposure 25 minutes. A very strong continuous spectrum extends from λ 4610 to 3440, continuing more weakly to 3057, *loc. cit.*, KAYSER and RUNGE.

Ivory scale-numbers.	$\frac{1}{\lambda}$	KAYSER and RUNGE.		Remarks.
		$\frac{1}{\lambda}$	λ	
21.1	1714	1714610	5832.23	S.s. } *
22.3	1723	1723541	5802.01	D.s. } B group
22.8	1729	1729305	5782.67	D.s. } L. DE B.
34.9	1866	1865713	5353.6	S.s. } Measured also by
35.4	1873	1872631	5340.08	S.s. } L. DE B.
96.15	2471	2470746	4047.36	P.s. } Measured also by
96.3	2473	2472622	4044.23	P.s. } L. DE B.
148.7	2902	(2900661) (2901503)	(3447.49) (3446.49)	P.s.
176.7	3110	31080	(3217.76) (3217.27)	P.s.

P.s. Principal series.

D.s. Diffuse series.

S.s. Sharp series.

CADMIUM.

Metal and also cadmium sulphate yield the same spectrum, consisting of one line only. It is the least refrangible of the triplets at Cd 17. Exposure 30 minutes.

Scale-numbers.	Oscillation frequencies from curve.	Oscillation frequencies for comparison.	Wave-lengths.	
170.9	30663	3066384 K. and R.	3261.17 K. and R.	KAYSER and RUNGE.

ZINC AND ZINC OXIDE.

Zinc foil was burnt in the oxy-hydrogen flame during 30 minutes. Nothing but a continuous spectrum is visible. Zinc oxide was intensely ignited in the flame for 60 minutes; it yielded nothing but a continuous spectrum. No lines or bands were visible.

* Measured also by LECOQ DE BOISBAUDRAN.

FLAME SPECTRA AT HIGH TEMPERATURES.

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CALCIUM FLUORIDE.

The substance used was fluor spar. Exposure 40 minutes.

Ivory scale-numbers.	$\frac{1}{\lambda}$	λ	Remarks.
12.5	16094	6213.5	The centre of a band
17.2	16642	6009	" " "
{ 24 to	17425	5739	A faint band
28.2	17910	5583.5	
{ 28.2 to	17910	5583.5	Band stronger than the preceding
30.5	18171	5503	
{ 35.15 to	18683	5352.5	Band
36.3	18812	5316	
{ 36.3	18812	5316	Band
36.7	18855	5303.5	
84.35	23637	4231	A strong line.

The last is possibly a line measured in the calcium spectrum by KAYSER and RUNGE λ 4226.91.

'Ueber die Spectren der Elemente,' Königl. Preuss. Akademie, 1891, IV.

STRONTIUM OXIDE.

Strontium sulphate was the substance used. Exposure 30 minutes.

Ivory scale numbers.	$\frac{1}{\lambda}$	$\frac{1}{\lambda}$	λ .	Remarks.
{ 15.4 to	16434		6085	A band.
16.3	16520		6053	
21.9				Weak line.
29.2	18028	1803918*	*5543.49	Weak nebulous line.
63.8	21697	2170365*	*4607.52	Strong line.
64.8	21780		4591	Faint. Sr?
84.25	23650	2365794*	*4226.91	Faint.
85.0	23700		4216.5	Faint.
93.9	24517	2452254*	*4077.88	

* Lines measured by KAYSER and RUNGE. 'Ueber die Spectren der Elemente,' Königl. Preuss. Akademie, 1891, IV.

BARIUM OXIDE.

Barium Sulphate. Exposure 30 minutes.

Ivory scale numbers.	$\frac{1}{\lambda}$	λ	Remarks.
{ 24.5 to	17483	5720	} The centre of a weak band.
	17508 mean	5712	
{ 25	17551	5697	} A strong band overlapping a weak one.
{ 25.3 to	17575	5690	
{ 26.1 continues to	17667	5660	
{ 27.2	17795	5619.5	} A band which is weakened between 28.9 and 30.9.
{ 28.0 to	17900	5587	
{ 28.9 continues to	18002	5555	} A line lies on the preceding P.s. band.
{ 30.5	18183	5499	
{ 29.3	18037	5544	} End of band, sharp.
{ 30.5	18170	5503	
{ 34.0 to	18572	5384	} End of band, sharp.
{ 35.0	18670	5356	
{ 36.0 to	18789	5322	} End of band, sharp.
{ 39.5	19154	5221	
{ 41.5 to	19373	5162	} Band.
{ 44.0	19648	5089.5	
{ 50 to about	20275	4932	} Stronger part of band.
{ 54			
{ 51.8 to	20463	4887	
{ 52.8	20565	4862.5	} Very faint band.
{ 54.0	20690	4833	
{ 59 to	21208	4715	}
{ 60	21312	4692	

CALCIUM OXIDE.

These measurements are taken from dolomite and from pure line.

Micrometer measurements in hundredths of an inch.	Description of the spectrum.	$\frac{1}{\lambda}$	λ .	Remarks.
11.33	} The less refrangible edge of a strong band gradually fading on its more refrangible side	15992	6253	} A strong Ca line in arc, 5594.64. K. and R.
14.76		16355	6116	
15.72				
20.0	} The more refrangible of the same } } The less refrangible edge of a weaker band, the more refrangible edge of which is coincident with the sodium line at 20, though the band is strong only as far as 18 }	16467	6075	
23.95		16969	5895	
27.72		17429	5739	
	} The less refrangible edge of a stronger band lying upon the foregoing upon, or is continuous with, the foregoing	17867	5598	
		18238	5485	
		18368	5445	
31.0	} The more refrangible edge of the same }	18445	5422	
32.17		18550	5390	
32.87		18665	5359	
33.8	} A marking in the same not very distinct }	18729	5341	
34.85		18796	5322	
35.42		18857	5304	
36.06	} The less refrangible edge } } This band is faint and becomes almost imperceptible when magnified ten diameters.			
36.64				
84.63	} The less refrangible edge of a very narrow band like a very strong broad line	23688	4222	} 4226.91 Ca line, r. in arc, very strong. K. and R.
85.2		23724	4215	

PHOSPHORUS PENTOXIDE.

A strong continuous spectrum extends from near the yellow sodium line to about λ 4090. A number of lines were observed many of which were identified with those of iron at wave-lengths 3888·2, 3860·5, 3749·4, 3747·6, 3736·9, 3733·5, 3722·8, 3720·2, 3705·5, 3440·2, and 3431·1. The following lines, all very faint, were not identified with any other substance, and it is probable that they are indications of a feeble band spectrum.

Ivory scale-numbers.	λ .
168·5	3279
169·2	3274
169·7	3271
170·05	3268
171·6	3255
172·9	3245

ARSENIC.

This element gave a faint nebulous line at 168·4 or λ 3280, which approximates the first line in the P_2O_5 spectrum.

SELENIUM.

Micrometer measurements in hundredths of an inch.	Description of the spectrum.	$\frac{1}{\lambda}$	λ .	Remarks.
51.68	The fainter edge of a band	20449.5	4890	4745, SALET (spark), 475 by combustion.
54.71	The stronger edge of the same.	20763.5	4816	
55.22	The fainter edge of a band	20817	4804	
57.66	The stronger edge of the same.	21070.5	4746	
58.78	The fainter edge of a band	21186	4720	4675 (PLÜCKER and HITTORF), spark. Also SALET.
60.73	The stronger edge of the same.	21387	4676	
62.20	The fainter edge of a band	21535	4643	4675 (PLÜCKER and HITTORF), spark. Also SALET.
64.28	The stronger edge of the same.	21744	4599	
65.70	The fainter edge of a band	21884.6	4569.5	
69.50	The stronger edge of the same.	22264.4	4491.5	
Here follow a series of 18 bands, which appear to overlap.				
73.78	The stronger edge of the 1st band	22688	4407.5	4675 (PLÜCKER and HITTORF), spark. Also SALET.
77.65	" 2nd "	23048	4339	
80.01	" 3rd "	23259.5	4299	
84.58	" 4th "	23683.5	4222	
87.78	" 5th "	23976.5	4170.5	
90.75	" 6th "	24249.2	4124	
92.99	" 7th "	24433	4093	
96.55	" 8th "	24745	4041	
101.12	" 9th "	25148.2	3976.5	
103.72	" 10th "	25371.5	3941.5	
105.35	" 11th "	25507.5	3921.5	
108.23	" 12th "	25751.5	3883	
110.75	" 13th "	25965	3851	
112.75	" 14th "	26129.5	3827	
115.29	" 15th "	26342.8	3796	
119.18	" 16th "	26672	3749	
120.55	" 17th "	26788	3733	
123.15	" 18th "	27002	3707	

TELLURIUM—(continued).

Micrometer measurements in hundredths of an inch.	Description of the spectrum.	$\frac{1}{\lambda}$.	λ .	Remarks.
91.01	The fainter edge of a band	24272	4120	4119.7. A line in spark spectrum, HARTLEY and ADENEY.
92.0	} Markings in the same } Darker edge of the same	24347	4107	
92.62		24400.5	4098	
93.505		24478	4085	
94.395		24555.6	4072.5	4072.7. A line in spark spectrum, HARTLEY and ADENEY.
97.67	The fainter edge of a band	24842.5	4025.6	
99.52	The stronger edge of the same.	25004.5	3999	
104.05	The darker edge of a band fainter than the foregoing	25398.5	3937	
108.51	" " " " " " " " " " " "	25775	3880	
117.54	" " " " " " " " " " " "	26533.5	3769	
122.68	" " " " " " " " " " " "	26964.5	3708.5	
127.05	" " " " " " " " " " " "	23717.6	3661	
132.29	" " " " " " " " " " " "	27744	3604	
136.70	" " " " " " " " " " " "	28091.5	3560	
155.95	A strong sharp line	29555	3383.5	3382.4. A line in spark spectrum, HARTLEY and ADENEY.
168.41	} A pair of lines } A strong sharp line, succeeded by the water vapour lines	30478.5	3281	3280. The same.
169.375		30550	3273	3273.4. The same.
172.635		30790.7	3248	3246.8. The same.

ANTIMONY.

A very good specimen of metallic antimony was used. Lead and copper were detected in it; the lines belonging to these elements being easily identified. Exposure, 30 minutes.

Micrometer measurements in hundredths of an inch.	Description of the spectrum.	$\frac{1}{\lambda}$.	λ .	Remarks.
30-155	The stronger edge of an indistinct band	18144	5511	
60-6	The fainter edge of band	21390	4675.5	
68-96	The stronger edge of the same.	22209.5	4503	
74-4	A faint line, or edge of a band	22746	4399.6	
81-57	"	23404	4273	
90-23	"	24202.5	4132	
93-9	The edge of a faint band	24518	4079	
95-85	"	24683	4051	
96-74	"	24761.5	4038.5	
100-108	A line or edge of an exceedingly faint band	25063	3990	
103-21	"	25311	3949.8	
104-19	The edge of a band	25410	3935.5	
105-95	"	25557	3913	
106-19	"	25578	3910	
107-71	"	25707	3890	
110-62	"	25954	3853	
113-87	"	26222	3813.5	
116-78	An indistinct line, or the edge of a band	26469.5	3778	
118-94	The edge of a band	26653	3751.9	
120-25	"	26764	3748.5	
123-47	"	27027	3700	
124-59	A pair of very faint though distinct lines	27125.5	3686.5	
125-65	The edge of an ill-defined band (faint)	27203.5	3676	
127-0	"	27314	3661	
128-65	"	27442	3664	
130-26	A faint line or indistinct edge of a faint band	27579	3626	
132-49	"	27760	3602	
135-37	"	27989	3573	

BISMUTH.

A beautifully crystallized specimen of the metal was used. Exposure 30 minutes. It was found to give lines belonging to lead, thallium, copper, and silver. These were easily identified.

Micrometer measurements in hundredths of an inch.	Description of spectrum.	$\frac{1}{\lambda}$.	λ .	Remarks.
22·3	A series of overlapping bands. The weaker or less refrangible edge of the 1st band } A line or narrow band lying upon the 1st band } The more refrangible or stronger edge of 1st band } The stronger or more refrangible edge of the 2nd band	17225	5805·5	The bands of this spectrum were not so sharply defined as those of lead and tin, and could not be so accurately measured. When two or more readings were not alike, the mean was taken. This band is very faint, as also are the succeeding ones, their degraded edges appearing as markings on a continuous spectrum which gradually fades away. This is very strong and broad, extending from 58·92 to 59·1.
24·65		17500	5714·2	
26·03		19172·2	5215·7	
29·06		18010·7	5549	
32·85	The same of 3rd band " 4th " " 5th " " 6th " The bands at this point are feeble and not distinct. Second series of band degraded towards the red	18443	5422	
35·63		18750	5333	
36·4		18832	5310·1	
37·0		18895·5	5292	
53·3	Feeble indication of 1st band.	20615	4850·5	
58·7	Very feeble indication of a band	21165	4724·5	
59·02	Centre of a broad line or more refrangible edge of a band	21211	4714·5	
59·5	A line upon a band	21247	4707	
60·03	The stronger or more refrangible edge of the 2nd band	21315·5	4691	
61·0	The same of 3rd band	21400	4672·8	
62·75	" 4th "	21590·5	4632	
65·11	" 5th "	21826	4582	
66·93	" 6th "	22007·5	4544	
68·27	" 7th "	22141	4516·5	
69·87	" 8th "	22301	4484	
72·0	" 9th "	22514	4441·5	
73·12	" 10th "	22625	4420	
74·25	" 11th "	22733	4399	
A very weak band extends from 74·9 to 75·16. The more refrangible edge is the stronger.				

FLAME SPECTRA AT HIGH TEMPERATURES.

BISMUTH—(continued).

Micrometer measurements in hundredths of an inch.	Description of spectrum.	$\frac{1}{\lambda}$.	λ .	Remarks.
75.16	22817	4382.5	In HARTLEY and ADENEY'S spark spectrum of bismuth. The same, 2982.9. The same. There is a silver spark line at 2901.6, HARTLEY and ADENEY. In HARTLEY and ADENEY'S spark spectrum of bismuth.
76 to 77.0	There are feebly visible flutings from 76 to 77	22865	4373.5	
76.09	The stronger or more refrangible edge of a band	22970	4353.5	
92.8	There is a continuous spectrum as far as 92.8.	22904.5	4366	
96.2	A line or marking on a band	23140	4321.2	
105.24	This band continues to 105.4.	23498	4255.5	
109.02	The more refrangible edge of a feeble indistinct band	23823.5	3872.5	
111.28	The same	26008	3845	
118.98	"	26653	3752	
127.35	"	27342.5	3652	
140.0	"	28350	3527.9	
141.0	Three feeble lines	28417	3517.9	
141.9	"	28475	3510.5	
198.7	A strong line coincident with a water-vapour line	32601	3067	
205.5	"	33070	3023.8	
210.1	33386	2992.2	
210.8	33441	2983.1	
219.8	34031	2937.5	
227.0	A group of weak lines	34483	2900.2	
227.8	34511	2897.2	

LEAD.

Assay lead was used. Exposure 40 minutes. Several lines attributed by MITSCHERLICH to lead oxide are evidently related to the bands described and measured. These are indicated by M.

Micrometer measurements in hundredths of an inch.	Description of the spectrum.	$\frac{l}{\lambda}$.	λ .	Remarks.
25·63	More refrangible edge of 1st band degraded towards the less refrangible rays	17621	5675	
27·065	More refrangible edge of 2nd band not degraded	17791	5620·5	5615, PbO, M.
28·045	Less " " 3rd band, broad	17903·5	5585	
31·7	More " " same not degraded, or very feebly towards the less refrangible rays	18315	5460	5460, PbO, M.
33·51	More refrangible edge of a feeble band, 4th band	18517·5	5400	
35·4	" " band which overlaps the foregoing band, 5th band	18726·5	5340	5328, PbO, M.
38·77	Less refrangible edge of 6th band not well defined	19079	5241	
39·86	More " " the same	19193	5210	5220, PbO, M.
40·04	Less " " 7th band	19252	5194	
42·26	More " " same	19456·5	5140	5144, PbO, M.
45·43	Marking of feeble band, 8th band	19797	5051	
48·11	Less refrangible edge of a band not well defined, 9th band	20078	4980·5	4993, PbO, M.
48·86	Marking on a feeble band, 10th band	20157	4961	
49·08	" " 11th "	20180	4955	
50·26	" " 12th "	20302	4925·5	
50·7	" " 13th "	20348	4914·5	4913, PbO, M.
51·23	" " 14th "	20402·5	4901·5	
51·43	" " 15th "	20424	4896	
53·23	More refrangible edge of a well-defined band, 16th band	20585	4858	4880, PbO, M.
54·375	" " " degraded towards the less refrangible rays, 17th band	20728·5	4824	4852, M. 4825, M.
57·56	Well defined, more refrangible edge of a band degraded towards the less refrangible rays, 18th band	21060·5	4748	
59·34	More refrangible edge of a band, in continuous spectrum, 19th band	21243·5	4707	
61·59	" " " " " 20th "	21474	4657	
63·85	More refrangible edge or marking of band, in continuous spectrum, 21st band	21702·5	4608	4664, PbO, M.

FLAME SPECTRA AT HIGH TEMPERATURES.

LEAD—(continued).

Micrometer measurements in hundredths of an inch.	Description of the spectrum.	$\frac{1}{\lambda}$.	λ .	Remarks.
64·35	More refrangible edge or marking of band, in continuous spectrum, 22nd band	21751	4597·5	4593, PbO, M.
68·65	More refrangible edge or marking of band, in continuous spectrum, 23rd band	22179	4508·5	
71·32	A line	22446	4455	4468, PbO, M.
75·84	A line or band marking (scarcely visible), in continuous spectrum, 24th band	22881	4370·5	4381, M.
79·13	More refrangible edge of a band, in continuous spectrum, degraded towards the less refrangible rays, 25th band	23177	4314·5	
84·39	More refrangible edge of a band, in continuous spectrum, degraded towards the less refrangible rays, 26th band	23666	4225·5	
88·28	More refrangible edge of a band, in continuous spectrum, degraded towards the less refrangible rays, 27th band	24022	4163	
89·69	Apparently the more refrangible edge of a band, in continuous spectrum, degraded towards the less refrangible rays, 28th band	24152·5	4140·5	
95·35	A line very strong, broad	24638·5	4059	$\left\{ \begin{array}{l} 4062\cdot5 \\ 4058\cdot5 \end{array} \right\}$ arc, LIVEING and DEWAR.
97·48	More refrangible edge of a band, in continuous spectrum, degraded towards the less refrangible rays, 29th band	24826	4028	
100·33	More refrangible edge of a band, degraded towards the less refrangible rays, 30th band	25077	3985	
102·79	More refrangible edge of a band, degraded towards the less refrangible rays, 31st band	25290	3954	
105·92	More refrangible edge of a band, degraded towards the less refrangible rays, 32nd band	25854·5	3913	
108·48	More refrangible edge of a band, degraded towards the less refrangible rays, 33rd band	25773	3880	
111·79	More refrangible edge of a band, degraded towards the less refrangible rays, 34th band	26050	3839	
114·52	More refrangible edge of a band, degraded towards the less refrangible rays, 35th band	26282	3805	
116·42	More refrangible edge of a band, degraded towards the less refrangible rays, but weaker, 36th band	26438	3783	
119·91	A line on the more refrangible edge of a band, degraded towards the less refrangible rays, 37th band	26734	3740·5	
122·05	Feeble edge of a band, 38th band	26914	3715·5	

LEAD — (continued).

Micrometer measurements in hundredths of an inch.	Description of the spectrum.	$\frac{1}{\lambda}$.	λ .	Remarks.
124·95	Very strong, broad, and well-defined line	271·46	3684	
126·07	Very feeble edge of a band, 39th band	27237	3671·5	
127·53	" 40th "	27357	3655	
129·0	Very strong, broad, and well-defined line	27477	3639·5	
131·73	Feeble edge of a band, degraded towards the less refrangible rays, 41st band	27699	3610	
133·31	Feeble, more refrangible edge of band, degraded towards the less refrangible rays, 42nd band	27825	3594	
134·45	Feeble, more refrangible edge of band, degraded towards the less refrangible rays, 43rd band	27836·5	3592·5	
135·53	Feeble, more refrangible edge of band, barely degraded towards the less refrangible rays, 44th band	28001·5	3571	
137·17	Feeble, more refrangible edge of band, barely degraded towards the less refrangible rays, 45th band	28127	3555	
142·85	Very feeble edge of band, not clearly defined, 46th band	28560	3501·5	
144·48	Well-defined edge of a band, degraded towards the less refrangible rays, 47th band	28685	3486	
148·76	Feeble edge of a band, 48th band	29012	3447	
150·45	" 49th "	29141	3431·5	
153·45	Well-defined edge of a band, degraded towards the less refrangible rays, 50th band	29367	3405	
157·82	Very feeble marking in band, 51st band	29694	3368	
159·62	Still more feeble marking in band, 52nd band	29828	3352·5	
160·59	Very feeble edge of band, not clearly defined, 53rd band	29898	3345	
163·57	" 54th band " degraded towards the less refrangible rays	30120	3320	
165·12	Very feeble edge of 55th band	30235	3307	
165·59	Imperfectly defined edge of a double band, 56th band	30269·5	3304	
170·5	Feeble, more refrangible edge, well defined, of a band, degraded towards the less refrangible rays, 57th band	30635	3264	
177·77	Well-defined edge of broad band, also commencement of water vapour lines, 58th band	31157	3209·5	
239·29	Very well defined, weak, but sharp line	(35294	2832·2)	

TIN.

The spectrum of the metal; a very fine series of 47 narrow bands extends from near the sodium line to wave-length 3033·1. Exposure 30 minutes. These bands are degraded towards the rays of least refrangibility.

Micrometer measurements in hundredths of an inch.	Description of the spectrum.	$\frac{1}{\lambda}$.	λ .	Remarks.
61·07	The stronger edge of a band first	21422	4668	4605 to 4595 SnO ₃ , SALET.
61·63	” ”	21478	4656	
63·79	Feeble edge of a band not well defined	21695·5	4609	
66·31	” ”	21945	4557	
67·54	” ”	22068	4532	
68·8	” ”	22195	4505·5	
71·37	” ”	22449	4456	
71·7	The more refrangible edge of a band well defined, degraded towards the rays of least refrangibility	22450·5	4454	
75·92	The more refrangible edge of a band well defined, degraded towards the rays of least refrangibility	22888	4369	
77·16	Feeble edge of a well-defined band	23003·8	4347	
79·68	” ”	23228	4305	4244 to 4236 SnO ₃ , SALET.
82·0	More refrangible edge of a band, degraded towards the rays of least refrangibility	23445	4265	
83·33	More refrangible edge of a band, but apparently not degraded ” degraded towards the rays of least refrangibility	23569·5	4243	
84·63	More refrangible edge of a band, degraded towards the rays of least refrangibility	23688	4221·5	
90·51	More refrangible edge of a band not clearly defined, degraded towards the rays of least refrangibility	24227	4128	
91·21	More refrangible edge of a band, degraded towards the rays of least refrangibility	24278	4119	
93·25	More refrangible edge of a band, degraded towards the rays of least refrangibility	24455·5	4089	
97·11	More refrangible edge of a band, degraded towards the rays of least refrangibility	24793·5	4033	
100·085	More refrangible edge of a band, degraded towards the rays of least refrangibility	25120	3981	
102·75	More refrangible edge of a band, degraded towards the rays of least refrangibility	25285	3955	4083 to 4077 SnO ₃ , SALET.
106·4	More refrangible edge of a band, degraded towards the rays of least refrangibility	25595	3907	
109·2	More refrangible edge of a band, degraded towards the rays of least refrangibility	25835	3871	

TIN—(continued).

Micrometer measurements in hundredths of an inch.	Description of the spectrum.	$\frac{1}{\lambda}$.	λ .	Remarks.
111·6	More refrangible edge of a band, degraded towards the rays of least refrangibility	26055	3841	
112·78	More refrangible edge of a band, and not clearly defined band	26132	3827	
114·14	"	26244	3810	
116·05	"	26423	3787	
118·18	Feeble, ill-defined, edge of a band	26588	3761	
121·04	"			
123·8	" strong, and well defined, more refrangible edge of a band, degraded towards the rays of least refrangibility	26830	3727	
131·0	Very strong, and well defined, more refrangible edge of a band, degraded towards the rays of least refrangibility	27054	3696	
133·71	The feeble edge of a band	27640	3618	
138·02	More refrangible edge of a strong band, degraded towards the rays of least refrangibility	27857	3590	
144·07	Feeble, ill-defined, more refrangible edge of, or marking in, a band	28190	3547	
148·32	Well defined, more refrangible edge of a strong band, degraded towards the least refrangible rays	28653	3490	
151·62	Ill-defined edge of a feeble band	28977·5	3451	
154·725	Well-defined edge of a band, degraded towards the rays of least refrangibility	29228	3421	
162·43	Well-defined edge of a band, degraded towards the rays of least refrangibility	29462·5	3394	
166·22	More refrangible edge of a very strong band, degraded towards the rays of least refrangibility	30036·5	3329·5	
170·08	Ill-defined edge of, or marking in, a feeble band	30316	3298·5	
174·34	More refrangible edge of a well-defined, strong band, degraded towards the rays of least refrangibility	30602	3268	
178·25	Well-defined, more refrangible edge of a band, degraded towards the rays of least refrangibility	30915	3234·5	
182·0	Well-defined, more refrangible edge of a stronger band than the foregoing, degraded towards the rays of least refrangibility	31190·5	3206	
194·41	Well-defined, more refrangible edge of a stronger but narrower band than the foregoing, degraded towards the rays of least refrangibility	31461	3179	
198·29	The well-defined edge of a band among the water-vapour lines	32310	3095	
203·07	"	32582	3068·6	
207·11	Very feeble band marking	32907	3038·8	
211·735	"	33186·5	3031	
	"	33454	2989	

SILVER.

Exposure 30 minutes.

Micrometer measurements in hundredths of an inch.	Description of spectrum.	$\frac{1}{\lambda}$.	λ .	Remarks.
28.85	Faint indication of a line or marking	17796.5	5556.7	5556.6, THALÉN.
30.05	"	18132.5	5515.0	
30.98	"	18236	5483.7	5486.6, THALÉN.
31.51	"	18303.6	5463.4	5464.1, THALÉN.
59.82	"	21294	4696.0	
61.875	Less refrangible edge of 1st band	21502.3	4650.8	There is a pair of lines in the green which do not appear on the photographs.
63.465	More "	21663	4616.5	
64.665	Less refrangible edge of 2nd band	21781.6	4591.0	
65.99	More "	21913.4	4563.4	
67.42	Less refrangible edge of 3rd band	22056.8	4533.4	
68.15	More "	22129.2	4519.0	
69.54	Less refrangible edge of 4th band	22268.2	4490.9	4518, L. DE B., in AgNO ₃ sol.
70.54	More "	22367.5	4470.9	4475.1, THALÉN, also L. DE B.
71.59	Less refrangible edge of 5th band	22473	4449.8	
72.87	More "	22600	4424.8	
73.73	Less refrangible edge of 6th band	22683.3	4408.6	
74.41	More "	22747	4396.2	
75.78	Less refrangible edge of 7th band	22865	4373.5	4396, L. DE B., in AgNO ₃ sol.
76.39	More "	22932	4360.7	
77.17	Less refrangible edge of 8th band	23004.6	4347.0	
78.12	More "	23090	4330.9	
79.935	Less refrangible edge of 9th band	23252.5	4300.5	
80.28	More "	23284.6	4294.7	
80.955	Less refrangible edge of 10th band	23347	4283.2	
81.5	More "	23398	4273.9	
82.435	"	23484.8	4258.0	
82.985	The 11th band is very indistinct, and is followed by a diffused spectrum extending from about 82.435 to 84.1, with markings measured at—	23536	4244.9	
83.64		23595	4238.2	
84.05		23635	4231.0	
	Markings			

SILVER—(continued).

Micrometer measurements in hundredths of an inch.	Description of spectrum.	$\frac{1}{\lambda}$.	λ .	Remarks.
87·435	} Another band commences at 87·435, extending to 86·685, and overlaps another fainter band, which extends to 89·25. } Three faint markings, like fine lines in a portion of continuous spectrum A fine sharp line Marking like an exceedingly faint sharp line " " " " " " " " " A distinct fine line " " " " } A pair of faint lines A strong group of 8 fine silver lines. These increase in strength with increase in refrangibility, and become closer together (commencing at 148·095).	23944·2	4176·4	3541·3, H. and A.
86·685		24059·4	4156·4	
89·25		24112	4147·4	
92·37		24379	4102·0	
93·12		24444·5	4091·0	
93·32		24462	4088·0	
97·38		24817	4030·3	
115·85		26347	3795·4	
117·0		26488	3775·2	
122·3		26934	3712·7	
126·0		27232	3672·0	
129·75		27537·2	3631·5	
135·02		27960	3576·5	
138·82		28253·4	3539·5	
141·03		28421·5	3518·4	
141·93	28489·4	3510·0		
148·095	1st line of first group. 2nd " 3rd " 4th " 5th " 6th " 7th " 8th " Another group of 10 fine silver lines, which increase in strength with their refrangibility, and become closer together.	28960·5	3453·0	
148·345		28979·3	3450·7	
148·545		28995	3448·0	
148·78		29013	3446·7	
148·973		29027	3445·1	
149·405		29060·7	3441·1	
149·905		29098·6	3436·5	
150·37		29135	3432·3	
151·09	1st line of second group 2nd "	29139	3431·8	
151·56		29224	3421·8	

FLAME SPECTRA AT HIGH TEMPERATURES.

SILVER—(continued).

Micrometer measurements in hundredths of an inch.	Description of spectrum.	$\frac{1}{\lambda}$.	λ .	Remarks.	
151·925	3rd line of second group	29252·5	3418·5	Very strong line here in arc and spark. The rays extending from this line, which is the maximum of intensity, are not continuous, but consist of an exquisite series of fine lines very close together and growing wider apart as they become less refrangible, down to 153.	
151·985	4th "	29256	3418·1		
152·34	5th "	29282	3415·1		
152·85	6th "	29321	3410·5		
153·25	7th "	29351·8	3407·0		
153·675	8th "	29383·7	3403·2		
153·985	9th "	29403·2	3401·0		
154·305	10th "	29431·4	3397·8		
A third group consisting of 19 fine silver lines commences at 154·625. These increase in strength as they extend farther into the region of the less refrangible rays.					
154·625	1st line of third group	29455	3395·0		
154·915	2nd "	29477	3392·5		
155·275	3rd "	29504	3389·4		
155·65	4th "	29532·3	3386·2		
155·95	5th "	29555	3383·5		
156·17	6th "	29571	3381·7		
156·57	7th "	29600·6	3378·4		
157·1	8th "	29633·5	3374·7		
157·42	9th "	29664	3371·1		
157·985	10th "	29706	3366·3		
158·49	11th "	29743·5	3362·2		
158·95	12th "	29766	3359·5		
158·95	13th "	29778	3358·2		
159·02	14th "	29783	3357·7		
159·5	15th "	29818·2	3354·8		
159·945	16th "	29851	3350·0		
160·405	17th "	29885·7	3347·2		
160·43	18th "	29887·5	3347·0		
160·95	19th "	29925·6	3341·6		
These groups of lines really constitute broad bands degraded on the side of least refrangibility.					

SILVER—(continued).

Micrometer measurements in hundredths of an inch.	Description of spectrum.	$\frac{1}{\lambda}$.	λ .	Remarks.
161.59	A fourth group of 6 fine silver lines, which are so faint and indistinct that only approximate measurements could be obtained.	29973.2	3336.4	
161.88		29995	3333.8	
162.155		30015	3331.7	
162.32		30027	3330.4	
162.55		30044.5	3328.4	
162.65		30051.5	3327.4	
163.58	This group of lines is succeeded by a series of bands extending from 163.58 to 168.27.	30121	3319.9	
164.15		30163	3315.3	
164.895	Strongest edges of the 9 narrow bands constituting this series. These are slightly degraded toward the less refrangible rays. This variation is more marked with those bands which are farthest from the red end of the spectrum.	30218.3	3309.2	
165.36		30252.2	3305.5	
166.39		30328.5	3297.3	
166.73		30353.5	3293.5	
167.39		30402.5	3289.2	
167.85		30437	3285.5	
168.27		30468	3282.1	
168.985		30521	3276.4	Very strong line occurs here in (centre), both arc and spark.
From 168.42 to 169.53		This series terminates in a very (centre) strong band degraded towards the less refrangible end of the spectrum.	30569.2	3271.3
169.65	30588		3269.3	
169.9				

IRON.

When investigating the spectrum of iron a number of materials were used, namely, pure metallic iron, tool steel, spiegel-eisen, ferro-manganese, silico-spiegel, and ferro chrome. Of the compounds of iron the following were taken: ferric oxide, ferrous sulphide, ferrous phosphate. Exposure from 15 to 35 minutes, generally 30 minutes. Pure iron and its compounds give spectra which are identical. Ferrous phosphate, however, yields a spectrum which contains a band due to phosphorus pentoxide, and a line also which is observed in this phosphorus compound, and in no other substance which, up to the present, I have photographed.

The metal and its compounds emit more or less strongly a series of bands lying between λ 5928 and 5537 which belong to iron. Steel also emits bands due to manganese, and the strong pair of lines of this metal.

The lines occurring in ferric oxide spectra are indicated, the description of the spectrum, and also those which are known to be prominent solar lines.

The spectrum was photographed from TURTON'S tool steel. R. means normal lines in ROWLAND'S map. K. and R. (KAYSER and RUNGE'S) measurements; r means that they observed the lines to be reversed in the arc, from which their measurements were made. C. CORNU'S measurements. W., MARSHALL WATTS.

Micrometer measurements in hundredths of an inch.	Description of the spectrum.	$\frac{1}{\lambda}$.	λ .	Remarks.
19.0	A faint band extending from 19 to 24, after which it darkens up to 25.3	16870	5927.7	5930.25. Strong line, K. and R.
24.0		17426	5738.5	5688 W. Splendid double line in Bessemer iron, Spiegel - eisen and MnO ₂ spectra.
25.3		17575	5689.8	
25.3	A dark band extending from 25.3 to 27.2, after which it gradually becomes fainter up to 34.1; within this latter portion there are bands or broad lines with their centres at 27.9 and 29.5	17575	5689.8	5544, Brightest edge of band. W. Manganese.
27.2		17795	5619.4	
27.9		17875	5594.3	
29.5		18060	5537.1	
34.1	In ferric oxide there is a band with a maximum of intensity about 25, extending to 28, and decreasing towards and as far as 35.	18570	5385	

IRON—(continued).

Micrometer measurements in hundredths of an inch.	Description of the spectrum.	$\frac{1}{\lambda}$.	λ .	Remarks.
36	Beyond the bands above-mentioned there is a continuation of diffused rays, or an indistinct band up to 155, in which are to be found the principal iron lines, all of which have been most carefully measured by LIVEING and DEWAR, and more recently by KATSER and RUNGE.	18780	5324.8	5324.31 R.
37.92		18988.5	5266.5	5266.72.
70.2		23325	4479.3	
71.2		22423	4459.7	4459.24.
73.0		22590	4426.7	4427.44.
74.2		22698	4405.7	4404.88. 4404 in Bessemer flame spectra, not identified, W.
75.4		22810	4384.0	4383.7.
75.8		22848	4376.8	4376.04 R.
78.7		23115	4326.2	4325.92 R.
79.8	FRAUNHOFER'S G	23210	4308.5	4307.96 R.
81.4	Two lines at this point measured as one	23406	4272.4	4271.93 } K. and R.
	A double line in appearance, but in reality a triplet			4271.3 } K. and R.
81.9		23436	4266.9	4267.97 R. K. and R.
94.4		24562	4071.5	4071.79.
95.37		24640.5	4058.3	
95.78		24676.7	4052.4	4052.75. K. and R.
96.13		24707	4047.5	4048.82 or 4045.9 R. K. and R.
97.22		24803	4031.7	4030.84. K. and R.
98.06	Strong double line highly characteristic of manganese	24876.5	4019.8	Manganese.
99.26		24982	4002.9	4005.9 ?
99.69		25020	3996.8	{ 3997.49 }
				{ 3998.16 }
100.83		25122	3980.6	
104.85	Observed in the ferric oxide spectrum	25465.5	3926.6	3928.05. K. and R.
105.3	" "	25503	3921.1	3923. K. and R.
105.72	" "	25538	3915.7	{ 3920.36 } K. and R.
				{ 3916.82 R. }
106.6		25613	3904.2	{ 3904.00 R. } K. and R.
				{ 3903.46 R. }

IRON—(continued).

Micrometer measurements in hundredths of an inch.	Description of the spectrum.	$\frac{1}{\lambda}$.	λ .	Remarks.
106.9	In Fe_2O_3 spectrum. There is a group of lines here which extends from $\lambda 3926.6$ to 3749.4	25638.5	3900.4	3899.8. K. and R.
107.09	•	25655	3897.8	3898.05. K. and R.
107.2	In Fe_2O_3 •	25664	3896.5	3895.75. K. and R.
107.34	•	25676	3894.6	3894.09. K. and R.
107.59	•	25697.5	3891.5	{ 3892.02. } K. and R. { 3890.96. }
107.9	In Fe_2O_3 •	25718	3888.2	3888.63. K. and R.
108.08	•	25739	3885.1	{ 3886.38. } K. and R. { 3885.61. }
108.47	•	25772	3880.2	
108.68	In Fe_2O_3 •	25789.7	3877.6	3878.82. K. and R.
108.93	•	25811	3874.3	3873.88 or 3872.61. K. and R.
110.0	In Fe_2O_3 •	25903	3860.5	3860.03. K. and R.
110.13	In Fe_2O_3 •	25914	3858.9	3858.49. K. and R.
110.55	•	25948	3853.7	3854.51. K. and R.
111.21	•	26005	3845.4	3853.7 C. 3846.96. K. and R. 3845.9 C.
111.53	•	26032	3841.4	3841.9. K. and R.
111.7	•	26047	3839.1	3840.58. K. and R.
112.15	•	26074	3835.2	{ 3836.48 3834.37. K. and R. 3827.96. K. and R. 3826.04. K. and R.
112.85	•	26137	3825.9	{ 3821.32. K. and R. 3821.32. K. and R.
113.1	•	26166	3821.5	
113.15	•	26170	3821.2	
113.25	•	26179.5	3819.7	3820.56. K. and R.
113.99	•	26242.5	3810.6	3810.89. K. and R.
114.18	•	26259	3808.1	3808.8. K. and R.
115.29	•	26342.5	3796.1	3975.13. K. and R.
116.18	•	26418.3	3785.2	3786.07. K. and R.
117.195	•	26504.8	3772.6	3773.84. K. and R.
117.7	•	26555	3765.3	3767.3. K. and R.
117.9	•	26572	3763.3	3763.9. K. and R.
118.3	In Fe_2O_3 •	26607	3757.9	3758.36. K. and R.
118.93	•	26652.5	3751.9	3751.97. K. and R.

IRON—(continued).

Micrometer measurements in hundredths of an inch.	Description of the spectrum.	$\frac{1}{\lambda}$.	λ .	Remarks.
119.1	In Fe ₂ O ₃	26675	3749.4	3749.61. K. and R.
119.25	"} A double line	26679	3748.1	3748.39. K. and R.
119.3	"}	26683	3747.6	3747.09 R. K. and R.
119.65	"} A prominent group of lines extending to λ 3705.5	26712.5	3743.5	{ 3745.67. K. and R.
120.1	In Fe ₂ O ₃	26755	3736.9	{ 3743.45. K. and R.
120.25	"	26770	3735.5	3737.27. K. and R.
120.96	"} A double line. CORNU'S M solar line	26823	3728.2	3735. K. and R.
121.0	"	26825	3727.9	3727.78. K. and R. M .
121.5	"	26865	3722.3	3727.13. K. and R.
121.7	"	26880	3720.2	3720.07. K. and R.
122.9	In Fe ₂ O ₃	26977	3705.5	3705.7. K. and R.
124.4	"	27103	3688.5	3687.77. K. and R.
124.9	"	27142	3685.8	3687.58. K. and R., or 3686.10.
125.15	In Fe ₂ O ₃	27162	3681.6	3680.43. K. and R.
128.2	"	27407	3648.6	3647.99. K. and R.
130.0	"	27550	3631.0	3631.62. K. and R.
131.9	"	27700	3609.2	3608.99. K. and R.
134.7	In Fe ₂ O ₃ . CORNU'S N solar line	27917	3581.1	3581.32. K. and R. N .
135.8	"	28005	3569.6	3570.23. K. and R.
136.1	"	28028	3565.0	3565.5. K. and R.
140.0	"	28328	3531.2	
143.1	"	28567	3501.8	3500.64. K. and R.
143.9	In Fe ₂ O ₃	28630	3492.3	3490.65. K. and R.
145.6	"	28760	3475.5	{ 3476.75. K. and R.
146.5	"	28835	3460.9	{ 3475.52. K. and R.
149.1	"} CORNU'S O solar line	29053	3440.8	{ 3460.02. K. and R. or 61.5 O .
149.4	"	29068	3440.2	3441.07. K. and R.
199.9	"	32689	3059.1	3440.69. K. and R.
201.8	"	32815	3047.4	3059.19. K. and R.
203.1	"	32905	3039.1	3047.71. K. and R. S .
205.9	In Fe ₂ O ₃ . CORNU'S T solar line	33100	3021.1	3040.54. K. and R. T .
	"		3021.15. K. and R.	{ 3020.7. K. and R.

NICKEL.

The metal and oxide were both examined, and the lines photographed were compared with those obtained by CORNU in the arc, and by LIVEING and DEWAR in the arc, the spark, in explosions of oxygen and hydrogen within tubes containing nickel, and also in the flame of nickel tetra-carbonyl.

CORNU "Spectre Normal du Soleil." 'Annales de l'École Normale,' 2 ser., vol. 9. 1880.

LIVEING and DEWAR, 'Phil. Trans.,' vol. 179, pp. 231-256, and 'Roy. Soc. Proc.,' vol. 52, p. 117.

The lines were measured by the ivory scale and were all identified with the exception of two, about which there is a slight doubt, namely 3574 and 3496.

The metal used was rolled nickel, which owes its malleability to a little manganese. The indications of the presence of this element were very evident from the bands between 5700 and 5300, and the double line 4031·8 and 4029·9.

The metal was exposed for half-an-hour, and the oxide, which yielded the better spectrum, one hour.

Ivory scale numbers.	$\frac{1}{\lambda}$.	λ .	LIVEING and DEWAR's measurements.	Remarks.
			λ .	
110·1	25900	3859	3857·8	Common to Ni(CO) ₄ , arc and spark spectra. Not seen in explosions. Unless exceptions are stated, all lines are common to the five different spectra (arc, spark, nickel carbonyl flame, and oxy-hydrogen explosions) as observed by LIVEING and DEWAR.
114·2	26255	3809	3806·6	
116·2	26426	3784	3783·0	
116·9	26485	3776	3775·0	
131·0	27628	3619	3618·8	
131·85	27695	3611	3609·8	
133·0	27785	3599	or 3612·1	
135·5	27980	3574	3597·0	
136·0	28020	3569	..	
140·25	28347	3527	3572·9, CORNU.	
141·25	28425	3518	3570·8, CORNU.	
141·8	28467	3513	3527·1	
142·8	28547	3503	3519·1	
143·6	28607	3496	3514·4	
144·5	28675	3487	3514·4	
145·8	28777	3475	3501·8	
147·0	28878	3462	3492·3	
147·25	28900	3460	3485·2	
148·0	28962	3453	3470·8	
148·75	29025	3445	3461·1	
149·75	29105	3436	3457·9	
150·1	29131	3433	3452·9	
151·2	29213	3423	3452·3	
152·2	29284	3415	3452·3	
154·8	29477	3392	3452·3	
155·0	29492	3391	3452·3	
156·1	29573	3381	3452·3	
157·4	29673	3370	3452·3	
163·9	30157	3316	3452·3	
174·5	30934	3233	3452·3	
			3445·7	Not in Ni(CO) ₄ nor explosions.
			3436·7	
			3433·0	
			3423·1	
			3413·8	
			3392·4	
			3390·4	
			3380·0	
			3371·3	
			3367·2	
			3368·9	} 3371·3 in Ni(CO) ₄ . Not in explosions. 3370 is probably the line 3368·9
			3315·1	
			3232·6	} Not in explosions.

COBALT.

The metal and oxide were both examined. The lines photographed were compared with those measured by LIVEING and DEWAR in the arc and spark. ('Phil. Trans.,' vol. 179, p. 231.)

Measurements were made with the ivory scale, and all the lines were identified.

The oxide and metal, as in the case of nickel, give the same spectrum. The exposure of the oxide was double that given to the metal. As in the preparation of malleable cobalt, some maganese is added; the bands and lines of this element appear in the photograph, but less distinctly than in the metallic nickel.

Scale numbers	$\frac{1}{\lambda}$.	λ .	LIVEING and DEWAR'S measurements.	Remarks.
			λ .	
91·2	24277	4119	..	4120 HUGGINS. On comparing the two series of wave-lengths it will be seen that the difference between them is rather larger than usual, which appears to be due to the scale not being quite accurately adjusted between certain points which are clearly indicated. The wave-lengths do not approximate so closely to LIVEING and DEWAR'S measurements as is the case with those in the nickel spectrum.
99·8	25026	3996	3997·3	
107·3	25640	3899	3905·2	
109·0	25804	3875	3873·2	
111·15	25990	3847·5	3844·8	
131·0	27628	3819·5	..	
131·85	27688	3612	3611·3	
132·6	27755	3603	3601·6	
133·25	27805	3596	3594·4	
135·1	27950	3578	3577·4	
			(more probably 3574·9)	
135·75	28000	3571	3568·9	
139·4	28280	3536	3532·8	
139·9	28320	3531	3529·3	
140·1	28335	3529	3528·4	
140·3	28352	3527	..	
141·4	28437	3517	3517·7	
141·8	28467	3513	3512·0	
142·15	28495	3509·5	{ 3509·3 } { 3509·7 }	
142·7	28537	3504	3502	
143·6	28607	3496	3495·1	
145·7	28710	3483	3482·7	
146·5	28835	3468	3465·2	
147·0	28878	3463	3462·2	
147·2	28895	3461	3460·6	
147·9	28953	3454	3452·9	
148·4	28995	3449	{ 3448·6 } { 3448·9 }	
149·0	29045	3443	3443·0	
150·2	29137	3432	{ 3432·9 } { 3432·4 }	
152·2	29285	3415	3414·2	
152·4	29300	3413	{ 3411·7 } { 3412·0 }	
152·9	29336	3409	3408·6	
153·3	29365	3405	3404·5	

CHROMIUM.

The spectrum obtained from ferro-chrome containing 22 per cent. of chromium, contains six lines due to chromium, and in addition bands and lines of iron. The bands extend from 24 to 28·3, and continue weaker as far as 35. Manganese lines are also very strong.

Ivory scale numbers.	Description of spectrum.	$\frac{1}{\lambda}$.	λ .	Remarks.			
80·6 81·4 82·75	A group of three well-defined lines	{ 2331 2338 2350	4290 4277 4255	ÅNGSTRÖM and THALÉN. 4289·4 4274·6 4253·9			
132·2 133·3 134·9					{ 27724 27810 27935	3607 3595 3580	LIVEING and DEWAR. 3606 3593 3578

CHROMIC TRIOXIDE.

This substance gives, in addition to the above, two groups of three lines, a continuous spectrum, strong, from close to the sodium line in the yellow, but a little less refrangible up to λ 3820.

IRIDIUM.

This element occasioned some difficulties. Strips of iridium, twisted into loops, were obtained from Messrs. JOHNSON and MATTHEY some years ago for the purpose of serving as supports for the alkalis and alkaline earths in the oxy-hydrogen blow-pipe. To this use it was put with some success and found convenient, but with oxides capable of undergoing reduction, even such as cupric oxide, it became corroded. It was found to be a convenient support for silicates which are fusible, but on examining the spectrum of silica, several lines were discovered which were not due to silica.

Three varieties of silica were tested—1st, Silica precipitated from sodium silicate. This yielded lines identified with iron even after treatment with hydrochloric acid. 2nd, Silica precipitated from silicon fluoride by passing the gas into water. The silica was evaporated from the hydrofluosilicic acid by filtration through absolutely pure ashless filter-paper. Even this showed a number of lines which at first were taken to be those of iron. 3rd, Rock crystal exposed to the hottest part of the flame on iridium for one hour gave nothing beyond the sodium lines in the yellow, mean λ 5892, and in the ultra-violet λ 3303.

To prove the origin of the lines which had been previously observed, a piece of

clean iridium was heated in the flame for seventy minutes and the spectrum photographed.

When the wire was at its highest temperature the flame assumed a peculiar bluish colour and the wire became very thin. The spectrum obtained proved to be similar to that previously obtained from pure silica.

A second spectrum was taken on the same plate, a little silica being placed on the loop of iridium. The spectrum was similar to the first, the lines being the same, but weaker, as the silica acted as a glaze and protected the wire.

It is perfectly evident that this metal was to some extent vaporized in the flame, and that the vapour emits a line spectrum.

The following are measurements of the lines photographed :—

λ .	λ .	λ .	λ .
4386		3599	3479
4256	3812	3596	3475
3965	3772	3533·5	3464
3937	3705	3511·5	3436·4
3860	3696	3508·7	3400
3815	3663	3484·3	3328

These lines have not yet been identified, but they are suspected to be due to osmium.

A small strip of pure iridium, for which I am indebted to Mr. GEORGE MATHEY, F.R.S., was exposed to the flame for three hours and a quarter, and a line spectrum with a small portion of a continuous spectrum was photographed. Undoubtedly the iridium was volatilized, for it lost weight to the extent of 0·0826 grm., and the end was worn away by the flame impinging upon it. The spectrum was very weak, the lines were not those referred to above, and it is suspected that some of them at least are due to a gaseous spectrum, or possibly to a series of the lines belonging to the spectrum attributed to water vapour which have not previously been observed.

The fact that iridium is slightly volatile has undoubtedly been proved, but if the metal is pure it may be used advantageously for the purpose of supporting irreducible oxides in the oxy-hydrogen blow-pipe flame.

ALUMINIUM.

When the metal, in the form of foil, is burnt in the oxy-hydrogen blow-pipe, it gives a spectrum which is continuous, but in which some few lines or narrow bands are visible. There can be little doubt that these are due to impurities, principally iron. With the exception of three the lines are all very faint. The measurements, which are only approximations, owing to the indefinite character of the lines, are the

following:— λ 4047 broad line, Fe; 4033, Fe; 4023, Fe; 4004.5 Fe; 3996, Fe; 3975, CaO; 3963?; 3947.5?; 3989?; 4013? The pure metal cannot be vaporized except by the arc and spark.* Evidence of this is afforded by the fact that an uncondensed spark yields a very beautiful band spectrum. The lines of silicon have been looked for but not detected in this spectrum.

Ivory scale numbers.	Description of spectrum.	$\frac{1}{\lambda}$.	λ .
6 20 120 170	} Continuous band of rays from 6 to 170 strong. Very intense from 20 to 120		
96.5 101.7 102.8		Lines at	24740 25300 25294

COPPER.

COPPER foil was heated in the flame. Two silver lines were observed in this spectrum, λ 3383.5 and 3282.1.

Micrometer measurements in hundredths of an inch.	Description of spectrum.	$\frac{1}{\lambda}$.	λ .	Remarks.
30.3	The centre of a broad line . .	18160	5506.5	This spectrum is partly due to CuO apparently. 3289.9, spark, HARTLEY and ADENEY. 3265.2, 3260.2, two spark lines, HARTLEY and ADENEY. The lines 3290 and 3262.5 are frequently seen in photographs where they would be least likely to be found.
44.35	A faint narrow line	19684	5080	
167.31	The centre of a broad line . .	30398	3290	
170.76	" " " . .	30652	3262.5	

* This statement is not quite correct. See Appendix (5).

COPPER OXIDE.

Micrometer measurements in hundredths of an inch.	Description of the spectrum.	$\frac{1}{\lambda}$	λ .	Remarks.
21-34	A faint line or narrow band, very indistinct	17125	5840	Feeble ray about 5858 in CuCl_2 spectrum, LECOQ DE BOISBAUDRAN.
22-61	The same, but a little stronger A continuous spectrum extends from λ 5747 to 4280, upon which are several bands superposed, the measurements of which are: The weaker and less refrangible edge of a fairly strong band The less refrangible edge of a strong band lying upon the foregoing	17272	5790	The following measurements are from LECOQ DE BOISBAUDRAN "Spectres Lumineux" Feeble ray about 5696.7
23-71	} There are indications of narrow dark bands at 28.8 and 30, overlapping the foregoing also at 32 and at 33.5	17402	5747	Band from 5584 to 5542, about middle 5563 CuCl_2 , 5545 CuO .
28-34		17939	5577	
28-8		18672	5356	
30-0		18880	5296	
32-0		19032	5241	
33-5	} The more refrangible edge of the same strong band The more refrangible edge of a narrow band overlapped by the foregoing	19298.5	5183	Centre of band about 5352
34-91		19587	5107	
36-86		20173	4954	
38-8	} The more refrangible edge of the broad strong line, which is coincident with the more refrangible edge of a weak band continuous with the foregoing bands, which are stronger	20922	4867	5239 in CuCl_2 5194.3 in CuCl_2 , 5195 in CuO 5106 Cu and CuO
40-82		20932	4777	
43-45		21225.5	4712	
49-2		21330	4688	
53-34	} The more refrangible edge of a narrower band overlapping the foregoing			4954 to 4938 CuO 4867 to 4847 CuCl_2 4777 CuCl_2 4704 approximately the less refrangible edge of a nebulous band, CuCl_2 4690 less refrangible edge of band indefinite and weak, CuCl_2
56-33				
59-16				
60-17				

FLAME SPECTRA AT HIGH TEMPERATURES.

COPPER OXIDE—(continued).

Micrometer measurements in hundredths of an inch.	Description of the spectrum.	$\frac{1}{\lambda}$.	λ .	Remarks.
62·16	The more refrangible edge of a narrow stronger band overlapping the foregoing	21531·5	4644	4642 CuCl ₂
68·2	The more refrangible edge of a broader and stronger band overlapping the foregoing	22134	4518	4522 to 4572 CuCl ₂
71·3	The more refrangible edge of a less diffuse band overlapping the foregoing	22444	4456	4453 approximately the middle of the maximum of light of a band, degraded, CuCl ₂
75·39	The more refrangible edge of a less well-defined weaker band	22839	4379	4369 maximum of light; this is variable. CuCl ₂
78·29	The more refrangible edge of a stronger band overlying the principal band	23105	4328	4330 to 4331 CuCl ₂
81·1	The more refrangible edge of a strong broad well-defined line, coincident with the more refrangible edge of the principal strong band extending to this point	23361·5	4280	4281 CuCl ₂
84·21	The more refrangible edge of a faint broad line coincident with the more refrangible edge of a fainter band, overlapped by the foregoing band	23650	4228	4217 about the more refrangible edge of narrow band of which the middle is at 4233
92·75	A very faint line, or faint marking	24412	4096	
93·92	"	24514·5	4080	
94·63	"	24576	4069	
95·72	"	24671	4053	
96·61	"	24750	4040	
97·29	"	24809	4031	
98·27	"	24895	4017	
168·18	The centre of a broad strong line	30461·5	3282	
171·5	"	30706	3256	

APPENDIX.

[1. Reference has been made to the fact that MITSCHERLICH ('POGG. Ann.,' vol. 121, p. 459, 1864) compares the band spectra of metalloïd elements with those of compound substances. He used both the oxy-hydrogen and oxy-coal-gas flames. He attributes only line spectra to copper, bismuth, lead, gold, iron, manganese, chromium, tin, potassium, sodium, lithium, zinc, cadmium, mercury, silver, barium, strontium, and calcium. He figures banded spectra of the following elements, magnesium, lines and bands; sulphur, selenium, tellurium, phosphorus, boron, iodine (bromine and chlorine, by absorption), and carbon.

Cyanogen and ammonia are also figured as giving channelled spectra, as well as the following metallic chlorides and oxides :—

PbO, PbCl₂, AuCl₃, Fe₂O₃ or FeO, MnO or Mn₂O₃, CuCl₂, CuBr₂, CuI₂, CuF₂, and CuO or Cu₂O, BiCl₃, BiBr₃, BiI₂, Bi₂O₃, BaO, SrO, CaO, BaF₂.

The following salts gave lines, or lines and bands together :—

BaCl₂, BaBr₂, BaI₂, CaF₂, CaCl₂, CaBr₂, CaI₂, SrF₂, SrCl₂, SrBr₂, SrI₂.

It will thus be seen that several metals enumerated on pp. 174 and 179 yield channelled emission spectra, and that these are not credited by MITSCHERLICH with other than line spectra, except in the case of magnesium, to which he assigns lines and bands. The most refrangible rays observed by MITSCHERLICH were about λ 4,000, and, though wave-lengths were not determined, the positions of lines and bands were measured and the spectra very carefully drawn.

2. LIVEING and DEWAR, in their "Investigations on the Spectrum of Magnesium," 'Roy. Soc. Proc.,' vol. 44, p. 243, give the following description of a spectrum ascribed to the oxide or to the process of oxidation :—

The component parts of the spectrum are the following—(1) The *b* group, λ 5183–5172–5166. (2) The MgH series, close to it, 5210, &c., and 5186, &c. (3) Bands in the green. (4) The triplet near L, λ 3838–3831–3829. (5) Triplet near M of the flame of burning magnesium, λ 3730–3724–3720, with the group of bands in that region. (6) The line, λ 2852.

The spectrum which I have described differs from the above inasmuch as the least refrangible ray photographed was λ 3929, which is at the edge of a strong band degraded towards the less refrangible side. Next, there is a strong line and a well-marked band, very strong from 3834 to 3805. LIVEING and DEWAR place the triplet near L, in or about this region. The triplet near M, and group of bands mentioned above, occupy the place of a band with lines upon it, extending on my photographs from λ 3805 to 3682.

Lines belonging to triplets near L and M were not recognized, though by varying the exposure and using sulphate, nitrate, and carbonate of magnesia, the conditions under which the spectra were obtained were modified. It is possible to obtain an intense continuous spectrum by prolonging the exposure to one hour and using the

nitrate. Strong lines are visible in the continuous spectrum or at its edge. LIVEING and DEWAR obtained their magnesia by burning the metal and holding the ash in the oxy-hydrogen flame.

The line λ 2852 is common to both spectra.

3. As to any possible relation of emission to absorption spectra, it may be remarked that ROSCOE and SCHUSTER found that there was apparently none in the case of sodium and potassium ('Roy. Soc. Proc.', vol. 22, p. 362, 1874), though the spectra were carefully measured.

LOCKYER and CHANDLER ROBERTS ('Roy. Soc. Proc.', vol. 23, p. 344, 1875) observed several channelled absorption spectra of metals by volatilizing them in tubes filled with hydrogen. No measurements were made, probably on account of the difficulties involved, and consequently the absorption spectra cannot be compared with channelled emission spectra of the same elements.

Channelled absorption spectra were observed in the vapours of silver, manganese, chromium, antimony, bismuth, and selenium.

Continuous absorption was noticed in copper, cadmium, iron, cobalt, nickel, tin, lead, gold, and palladium.

4. The spectra of sulphur, selenium, and tellurium were carefully investigated by SALET ('Ann. Chim. Phys.' [4], vol. 28, p. 47, 1873; also 'Traité Élémentaire de Spectroscopie,' p. 221), but only so far as the visible region, chiefly the green and blue rays. There are, however, many bands in the spectra of selenium and tellurium, which lie in the ultra-violet region, which appear on my photographs and have been measured.

5. According to a recent photograph, aluminium foil, when burnt, yields a beautiful channelled spectrum.

I have to express my thanks to Mr. HUGH RAMAGE, F.I.C., Assistant Chemist, Royal College of Science, for the care with which he has photographed many of these spectra, and otherwise rendered me valuable assistance.

W. N. H., *Sept.* 29, 1893.]

DESCRIPTION OF PLATES.

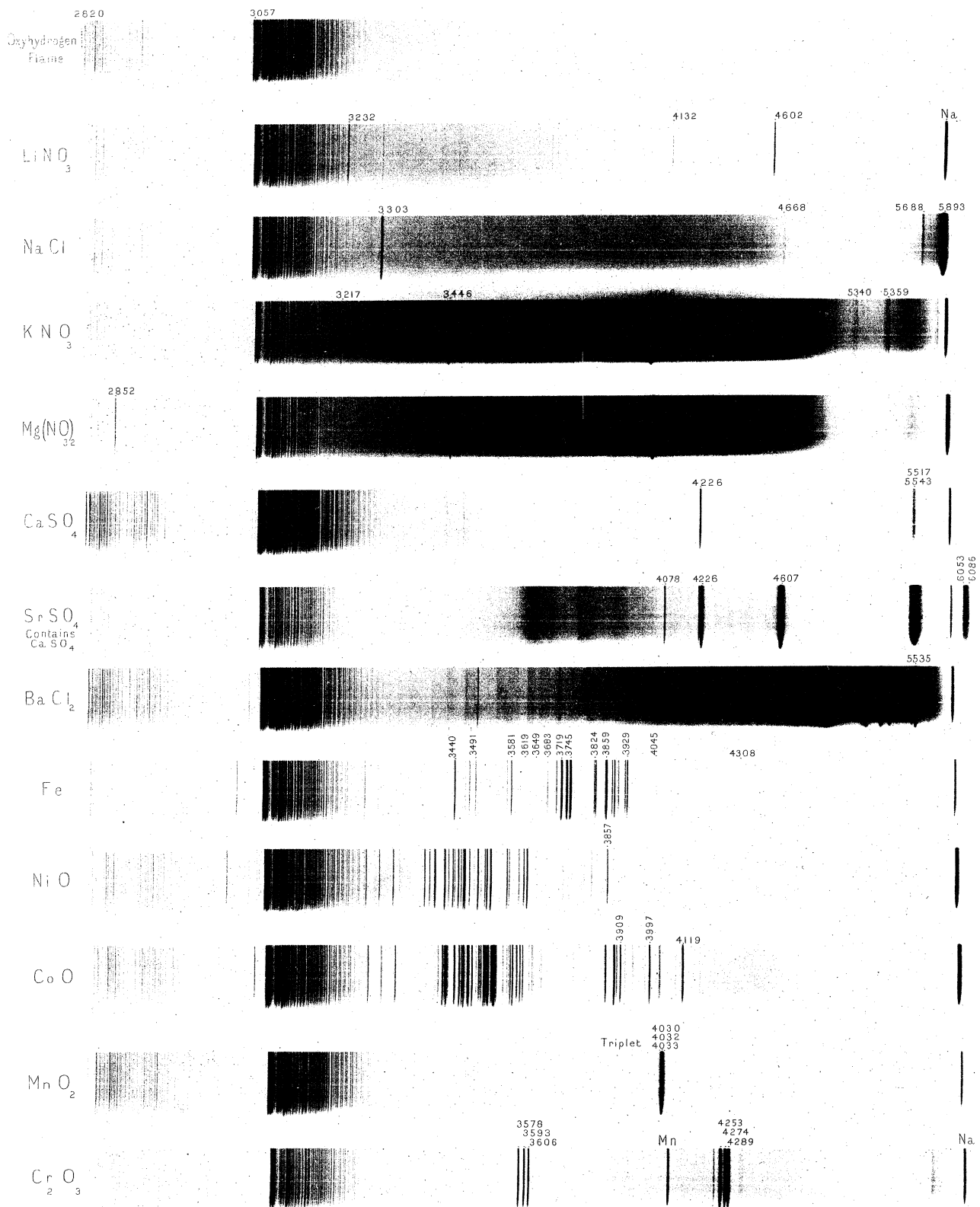
Photographs of spectra of the oxy-hydrogen flame, and of various salts, oxides, and metals, heated in the same for a uniform period of one hour. Dispersion used equal to one quartz prism of 60° . Enlarged about two diameters.

PLATE 6.

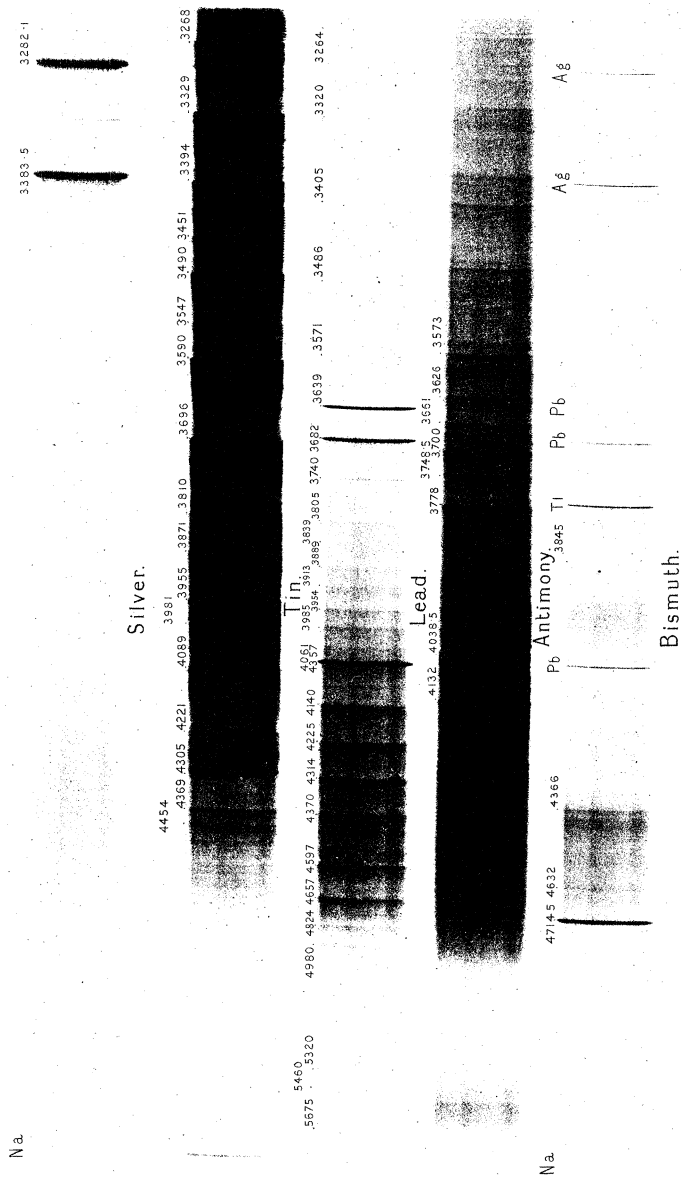
- | | |
|--|---|
| 1. Oxy-hydrogen flame, water vapour lines. | |
| 2. Lithium nitrate, lines of lithium and sodium. | |
| 3. Sodium chloride | } Band spectra of oxides and chlorides,
with line spectra of metals. |
| 4. Potassium nitrate | |
| 5. Magnesium nitrate | |
| 6. Calcium sulphate | |
| 7. Strontium sulphate | |
| 8. Barium chloride | |
| 9. Iron | } Line spectra of the metals chiefly. |
| 10. Nickel oxide | |
| 11. Cobalt oxide | |
| 12. Manganic oxide | |
| 13. Chromium sesqui-oxide | |

PLATE 7.

Band spectra of arsenic, antimony, bismuth, lead, and silver, with a dispersion of four quartz prisms of 60° . Enlarged about two diameters.

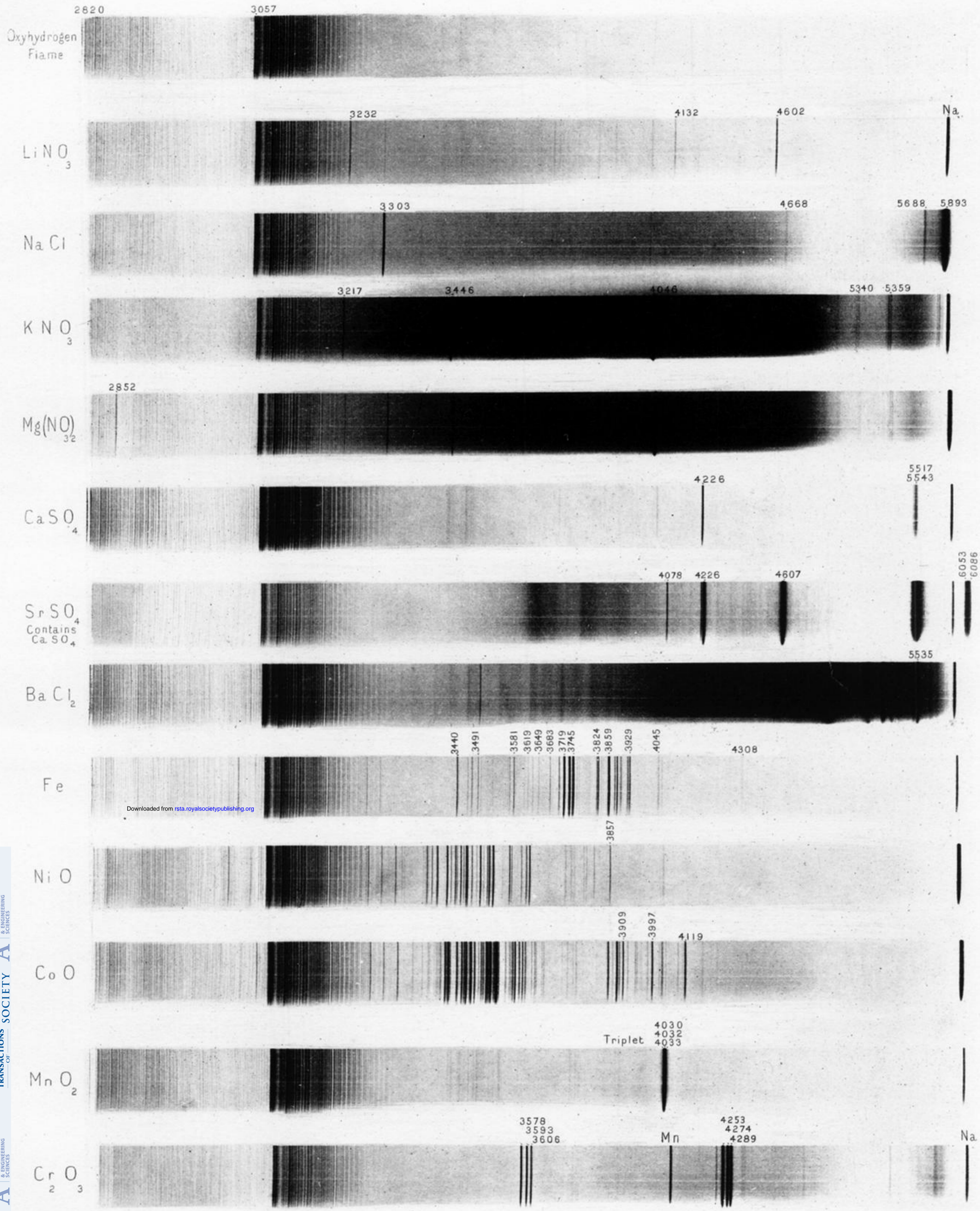


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Band Spectra of Metals.



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Na

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3383.5

3282.1

Silver.

4454

3981

4369 4305 4221

4089

3955

3871 3810

3696

3590 3547

3490 3451

3394

3329

3268

Tin.

5675 5460 5320

4980 4824 4657 4597

4370 4314 4225 4140

4061 4357

3985

3913

3839

3805

3740

3682

3639

3571

3486

3405

3320

3264

Lead.

4132 4038.5

3778

3748.5 3700

3661

3626 3573

Antimony.

4714.5 4632

4366

Pb

3845 Tl

Pb Pb

Ag

Ag

Bismuth.

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Band Spectra of Metals.